# Neutrinoless Double Beta Decay with <sup>136</sup>Xe

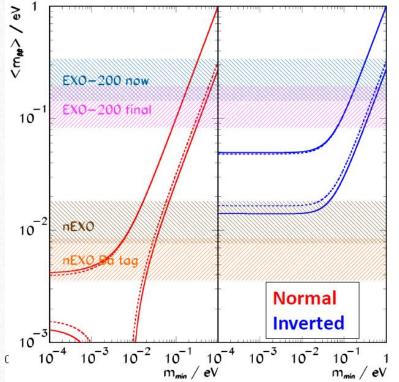
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**BNO-5**0

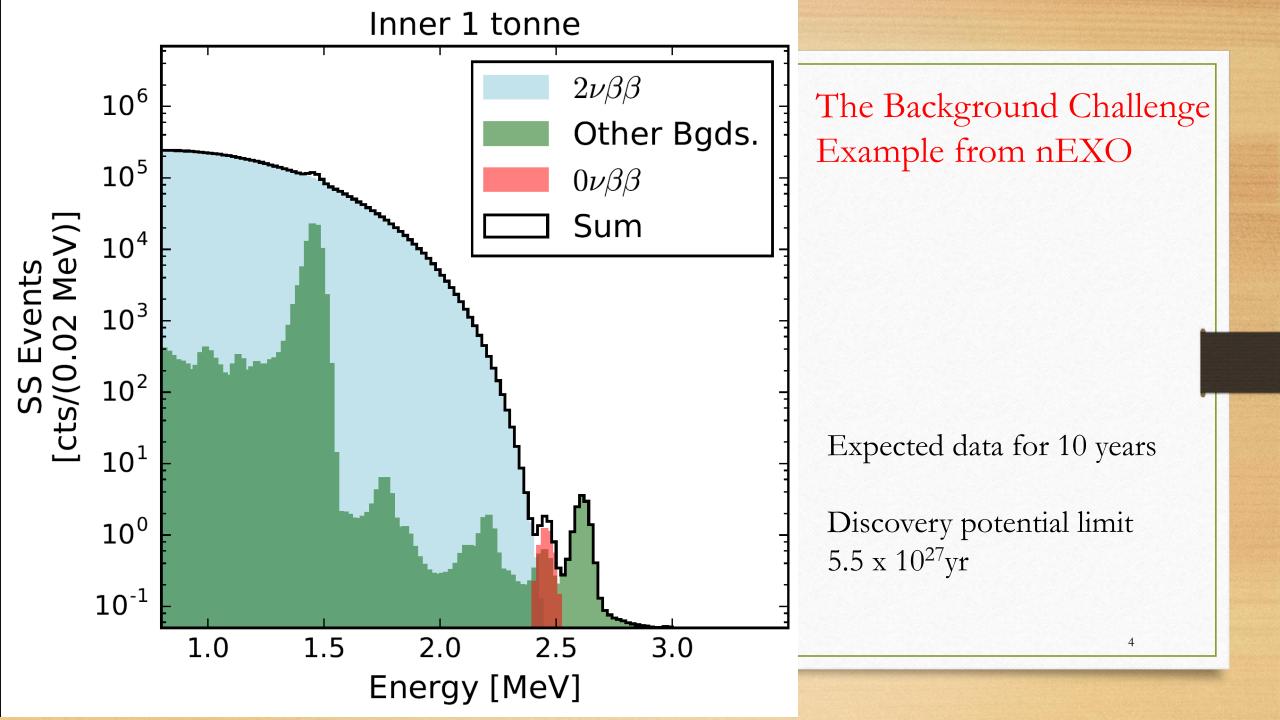
## The Challenge for Double Beta Decay

- Hints that data prefer Normal Hierarchy
- This suggests that we need to start thinking about how to detect NH
- Will require 100 fold reduction in backgrounds
- Will require 100 fold increase in ton-years
- I will speculate on how to get there!

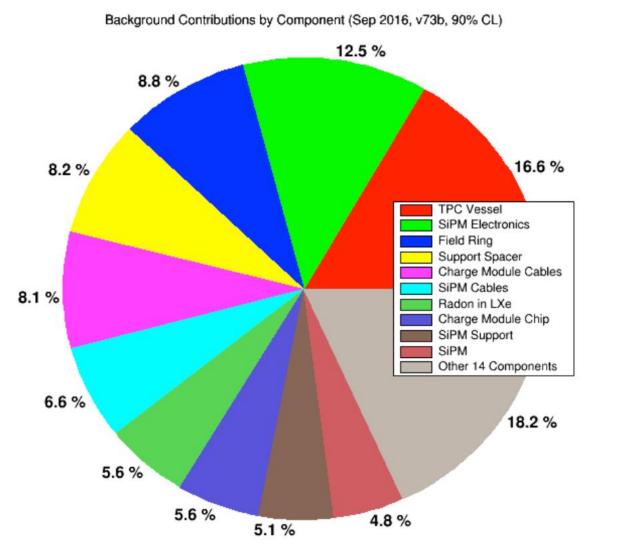


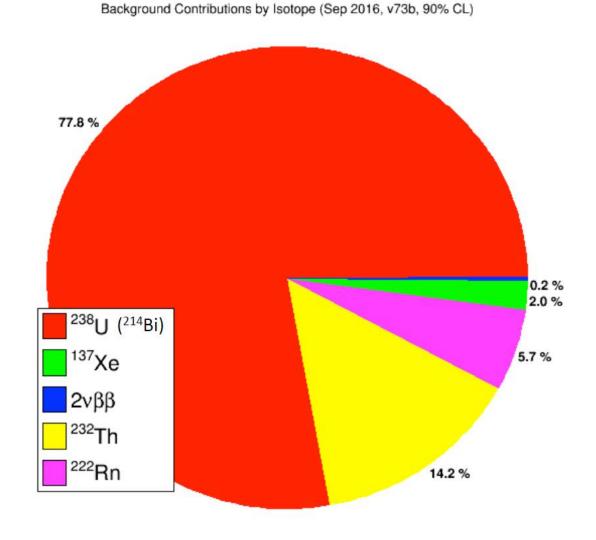
## Advantages for xenon

- Xenon can be made very clean
- Any cosmogenic activity can be removed by in situ purification
- No long lived radioactive isotopes (other than double beta decay)
- Least expensive of the double beta candidates to enrich
- Can be made into large, homogeneous detectors
- Concepts for detecting Ba daughter being developed



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## Possible detector configuration

- Dissolve in liquid scintillator (Kamland-Zen Yuri Efremenko's talk)
- Use in pure liquid form (EXO Vladimir Belov's talk)
- Use in pure, gaseous form (Goddhard, NEXT, Panda-X3)

## Dissolve in liquid scintillator

- Advantages:
  - Excellent shielding and self shielding
  - Excellent calorimetry
  - Modest energy resolution
- Disadvantages
  - Resolution limiting as one goes to normal hierarchy
  - Large target mass for solar neutrinos

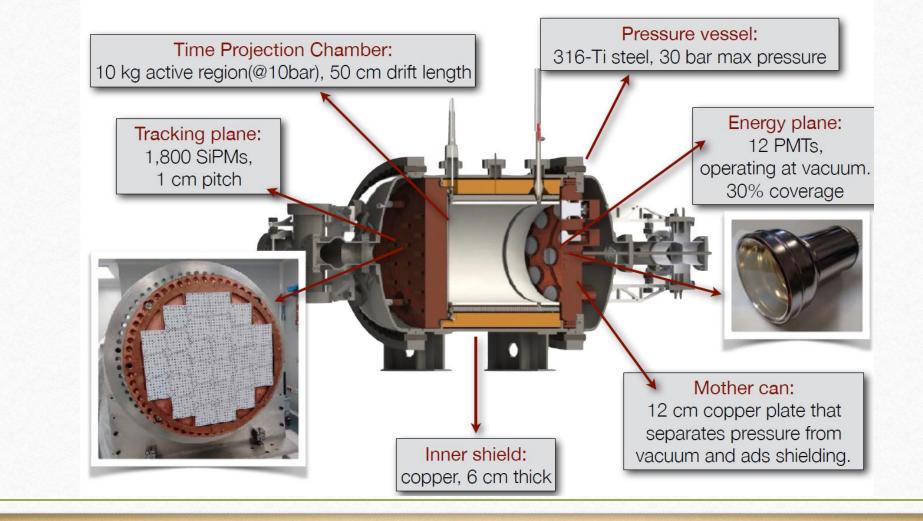
## Pure Liquid Xenon

- Good energy resolution (~1% sigma)
- Modest topological information
  - Important for background identification and rejection
- Good self shielding

#### Gas Detectors

- Excellent energy resolution (0.5% FWHM)
- Excellent topological information
- Very large -> large surface and hence backgrounds
- Probably best option for Ba tagging but long way to go

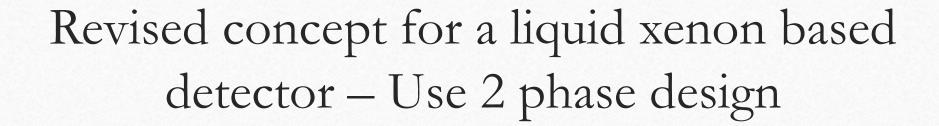
### NEW (NEXT-WHITE) at glance

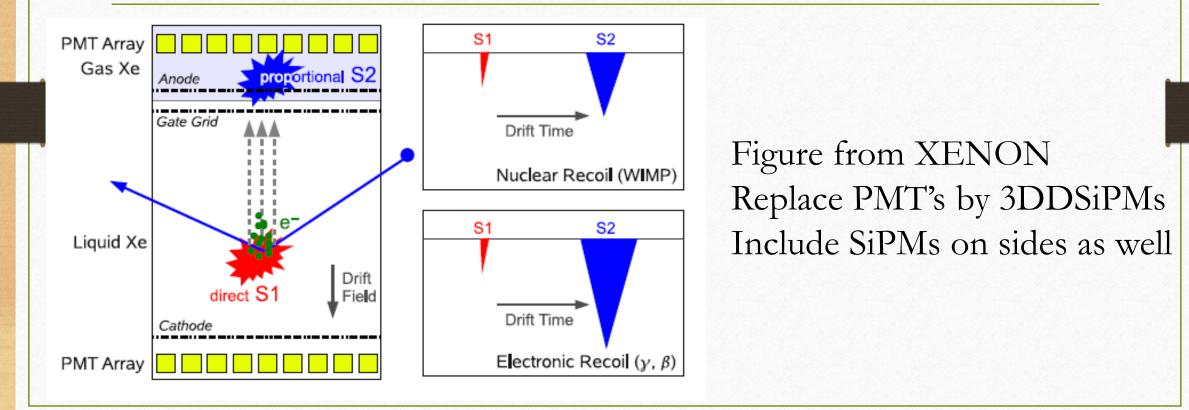


#### Data from NEXT Double electrons! (Co-56 data) Inclose\_weight uzu \_weigin hAnode\_weight hXY\_weight Entries 146 29867 Entries <sup>200</sup> = Mean x 42.6 0 200 Mean x 25.44 Mean y -53.93 Mean y -47.19 RMS x RMS y 150 \_\_\_\_\_ 100 \_\_\_\_\_ 50 \_\_\_\_ 0 \_\_\_\_ -50 \_\_\_\_ -100 \_\_\_\_ -150 \_\_\_ RMS x 43.74 60.2 150 RMS y 33.41 43.56 100 F 0.12 50 600 0.1 X-Y 0.08 -50 400 -100 E -0.06 -150 200 0.01 \_ -200 Ender der der bergen 50<sup>''</sup> 100 150 200 0.02 100 150 1. ... 1 I I I I 200 -200 -150 -100 -50 hXZ weight hYZ\_weight nXZ\_weight ntries 29857 hYZ\_weight 220 =220 F 29857 -53.12 148.5 Entries Entries - Mean x Mean y -41.73 148.5 Mean x Mean y 200 -200 33.12 25.77 RMS x RMS y RMS x RMS y 29.24 25.77 Ξ 0.25 180 ----180 \_ - 0.25 0.2 160 — 160 \_ Y-Z \_ -0.2 140 ----140 0.15 X-Z 0.15 120 120 0.1 0.1 \_ 100 \_\_\_\_ 100 0.05 0.05

## Importance of Topology

- EXO-200 gets gamma identification and some rejection based on single/multi site events ie almost all gammas Compton scatter
- Limitation Spatial resolution and high energy threshold limit effectiveness of the rejection
- In Gotthard experiment a factor of ~25 rejection was obtained by requiring 2 Bragg peaks
- In liquid topology of events is same but scale reduced  $\sim 50$  times
- Better resolution and lower threshold would also improve rejection
- Overall factor of 100 may not be impossible





## 3DDSiPM

- 3 dimensional, digital SiPMs being developed by Sherbrooke University in Canada
- Each SiPM pixel is treated as a logic bit, 0 or 1
- On chip processing to get, for example number of hits in given time in given area
- No analogue processing so no problem with capacitance in large arrays
- Very little heat (important as devices are in the cold xenon)
- Full area sensitive as logic is behind the SiPM layer
- Devices are working with  $\sim 100$  ps resolution

### More on SiPMs

- At -100 C, noise levels are very low (<<1 hit in a cluster signal)
- When combined with electroluminescence, noise  $\sim 10^4$  lower than pads
- Thresholds of 1 ionization electron possible

### How to use this

- In general in a liquid xenon detector, diffusion ~ few mm washes out the 2 peak structure
- With very fast and low noise signals we can used deconvolution to extract the original signal. Resolutions of few 10's µm possible for 1 metre detector (on paper) (use alphas for point spread function)
- Peak separations are ~ 1 mm
- Much better identification of gamma clusters possible
- May also be possible to catch the Cerenkov peaks from the initial electrons

## The way forward

- Prototypes being commissioned to see if the on-paper resolutions possible in practice
- Simulations will follow with realistic responses
- Also work in progress to assess economics of xenon isotope production by distillation (needs ~ 1 km high still – may be possible in mine shaft)

## Conclusions

- We may find double beta decay before the 1 meV scaleis a
- Getting to the full normal hierarchy will be very hard but perhaps not impossible
- Xe is a very attractive target
- We definitely need to keep our underground labs going for another 50 years