Determination of the  $\theta_{23}$  octant within and beyond the Standard Model

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#### NuFIT 3.0 (2016)

	Normal Ordering (best fit)		Inverted Ordering $(\Delta \chi^2 = 0.83)$		Any Ordering
	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	$3\sigma$ range
$\sin^2 \theta_{12}$	$0.306^{+0.012}_{-0.012}$	$0.271 \rightarrow 0.345$	$0.306^{+0.012}_{-0.012}$	$0.271 \rightarrow 0.345$	$0.271 \rightarrow 0.345$
$\theta_{12}/^{\circ}$	$33.56^{+0.77}_{-0.75}$	$31.38 \rightarrow 35.99$	$33.56^{+0.77}_{-0.75}$	$31.38 \rightarrow 35.99$	$31.38 \rightarrow 35.99$
$\sin^2 \theta_{23}$	$0.441^{+0.027}_{-0.021}$	$0.385 \rightarrow 0.635$	$0.587^{+0.020}_{-0.024}$	$0.393 \rightarrow 0.640$	$0.385 \rightarrow 0.638$
$\theta_{23}/^{\circ}$	$41.6^{+1.5}_{-1.2}$	$38.4 \rightarrow 52.8$	$50.0^{+1.1}_{-1.4}$	$38.8 \rightarrow 53.1$	$38.4 \rightarrow 53.0$
$\sin^2\theta_{13}$	$0.02166\substack{+0.00075\\-0.00075}$	$0.01934 \to 0.02392$	$0.02179^{+0.00076}_{-0.00076}$	$0.01953 \rightarrow 0.02408$	$0.01934 \to 0.02397$
$\theta_{13}/^{\circ}$	$8.46^{+0.15}_{-0.15}$	$7.99 \rightarrow 8.90$	$8.49^{+0.15}_{-0.15}$	$8.03 \rightarrow 8.93$	$7.99 \rightarrow 8.91$
$\delta_{\mathrm{CP}}/^{\circ}$	$261^{+51}_{-59}$	$0 \rightarrow 360$	$277^{+40}_{-46}$	$145 \to 391$	$0 \rightarrow 360$
$\frac{\Delta m^2_{21}}{10^{-5}~{\rm eV^2}}$	$7.50\substack{+0.19 \\ -0.17}$	$7.03 \rightarrow 8.09$	$7.50^{+0.19}_{-0.17}$	$7.03 \rightarrow 8.09$	$7.03 \rightarrow 8.09$
$\frac{\Delta m^2_{3\ell}}{10^{-3}~{\rm eV}^2}$	$+2.524^{+0.039}_{-0.040}$	$+2.407 \rightarrow +2.643$	$-2.514_{-0.041}^{+0.038}$	$-2.635 \rightarrow -2.399$	$ \begin{bmatrix} +2.407 \rightarrow +2.643 \\ -2.629 \rightarrow -2.405 \end{bmatrix} $

#### Some open questions

- What is the order of neutrino masses?
- Is there CP violation among neutrinos?
- Which octant does  $\theta_{23}$  belong to?

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#### Some other open questions

- Why are neutrino masses so small?
- Why are CP and P violated?
- Why do we have anomalies in the reactor, gallium and short-baseline data?

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..... beyond Standard Model physics is needed!

#### Beyond the Standard Model

Most attempts to find answer to the origin of the neutrino masses involve:

- Sterile neutrinos
- Non-standard interactions
- Majorana nature of neutrino

In case of one sterile neutrino

$$U_{4\times4} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

the  $3\times3$  matrix is no longer unitary.

#### Nonunitary mixing

• A convenient way to parameterize nonunitarity:

$$N = \begin{pmatrix} \alpha_{11} & 0 & 0 \\ \alpha_{21} & \alpha_{22} & 0 \\ \alpha_{31} & \alpha_{32} & \alpha_{33} \end{pmatrix} \times U_{PMNS}$$
$$H = \frac{1}{2E_{\nu}} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix}$$
$$+ N^{\dagger} \times \begin{pmatrix} V_{CC} + V_{NC} & 0 & 0 \\ 0 & V_{NC} & 0 \\ 0 & 0 & V_{NC} \end{pmatrix} \times N$$

 F.J. Escrihuela, D.V. Forero, O.G. Miranda, M. Tórtola and J.W.F. Valle, Phys. Rev. **D92**, 053009 (2015) and arXiv:1612.07377.

# Nonunitary mixing bounds: (Example from Escrihuela et al.)

Nonunitary parameter	Bound at 90% C.L.
α <sub>11</sub>	0.9974
α <sub>22</sub>	0.9994
α <sub>33</sub>	0.9988
<i>α</i> <sub>21</sub>	$2.6 imes10^{-2}$
\alpha_{31}	$2.0 imes10^{-3}$
$ \alpha_{32} $	$1.5 imes10^{-2}$

#### Nonunitary mixing

• Another way to parameterize nonunitarity:

$$\begin{aligned} H &= \frac{1}{2E_{\nu}} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} \\ &+ \frac{V_{CC}}{2} U^{\dagger} \times \begin{pmatrix} 2 - 2\alpha_{ee} & \alpha_{\mu e}^* & \alpha_{\tau e}^* \\ \alpha_{\mu e} & 2\alpha_{\mu\mu} & \alpha_{\tau\mu}^* \\ \alpha_{\tau e} & \alpha_{\tau\mu} & 2\alpha_{\tau\tau} \end{pmatrix} \times U \end{aligned}$$

- M. Blennow, P. Coloma, E. Fernandez-Martinez, J. Hernandez-Garcia and J. Lopez-Pavon, J. High Energ. Phys. 04, 153 (2017).
  - Both notations are equivalent when calculating the Hamiltonian and oscillation probabilities.

# Standard Model case (3 active neutrinos and nothing else)



# Nonunitary mixing (3 active and 3 sterile neutrinos, Blennow et al. bound)



# Nonunitary mixing (3 active and 3 sterile neutrinos, Escrihuela et al. bound)



# Light sterile neutrino (3 active and 3 sterile neutrinos, $0.1 \text{ eV}^2 < \Delta m_{41}^2 < 1 \text{ eV}^2$ , Blennow et al. bound)



# Light sterile neutrino (3 active and 3 sterile neutrinos, $\Delta m_{41}^2 > 100 \text{ eV}^2$ , Blennow et al. bound)



# Nonunitary mixing (Unconstrained $\alpha$ s)



## $|\alpha_{21}|$ dependency plot

