



# Development of particle identification systems based on aerogel at BINP

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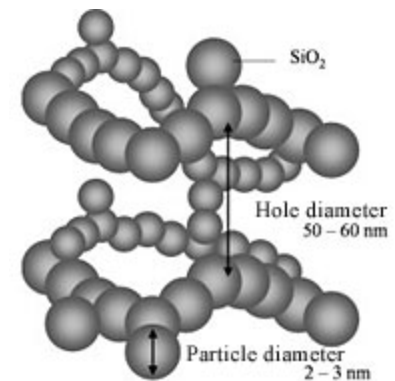
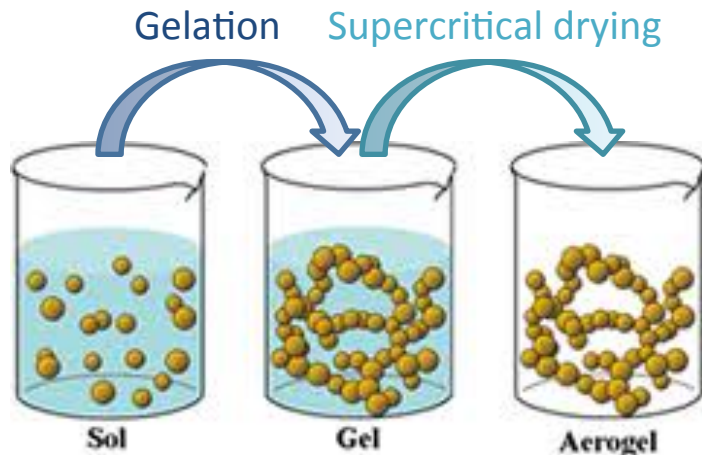
# Outline:

- Aerogel
- Threshold counters
  - ASHIPH method
  - The ASHIPH system of the KEDR detector
  - The ASHIPH system of the SND detector
- RICH detectors
  - AMS-02
  - CLAS-12 (J-Lab)
- Projects
  - FRICH for PANDA detector
  - FARICH for Super CTF
- Summary

# Silica aerogel



- Silica aerogel was first produced in 1931 by Samuel S. Kistler
- Lightest solids. Close the nature's gap in refractive index between gases @ STP ( $n-1 \lesssim 10^{-3}$ ) and liquids/solids ( $n \gtrsim 1.3$ ).
- 3D network of  $\text{SiO}_2$  nanometer sized pellets and 50-100 nm pores
- Now produced by sol-gel method out of silicon alkoxide  $\text{Si}(\text{OR})_4$





# Production of aerogel in Novosibirsk

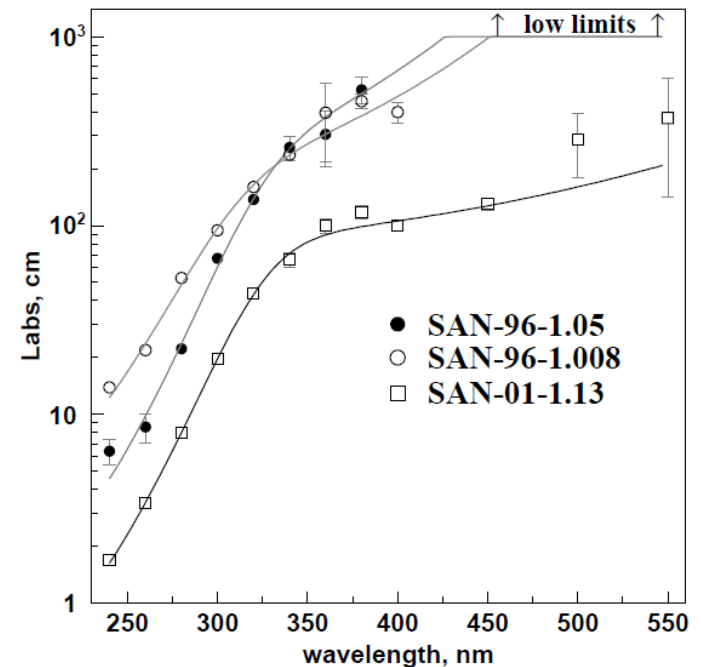
- Started in 1986 by the Boreskov Institute of Catalysis SB RAS in cooperation with the Budker Institute of Nuclear Physics SB RAS
- Hydrophilic
- Refraction indices **1.006 – 1.08** (1.08-1.13 produced by sintering)
- Block dimensions up to **200x200x50 mm<sup>3</sup>** (n=1.03) & **200x200x30 mm<sup>3</sup>** (n=1.05)
- Inner surface 800 m<sup>2</sup>/g
- Remarkable optical quality has been achieved:

$L_{abs}(400\text{nm}) = 5 - 7 \text{ m}$

$L_{sc}(400\text{nm}) = 4 - 6 \text{ cm}$

(Clarity = 0.0043 – 0.0064  $\mu\text{m}^4/\text{cm}$ )

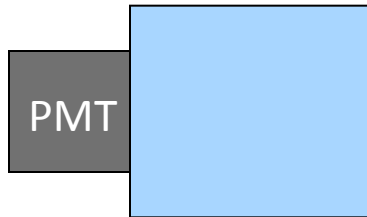
Absorption length vs  $\lambda$





# Threshold Cherenkov counters

## Direct light collection

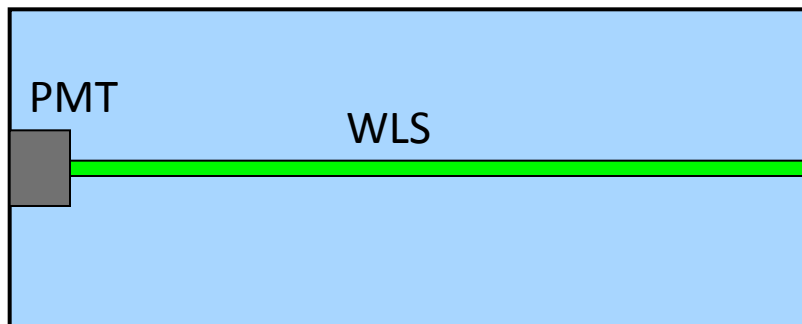


Pros: Simplicity

Cons: Counter size limited → large PMT number&area  
→ high total cost

## ASHIPH – Aerogel-**S**hifter-**P**hotomultiplier

Suggested by A.Onuchin et al. for PID of the KEDR detector [ [NIM A315 \(1992\) 517](#) ]



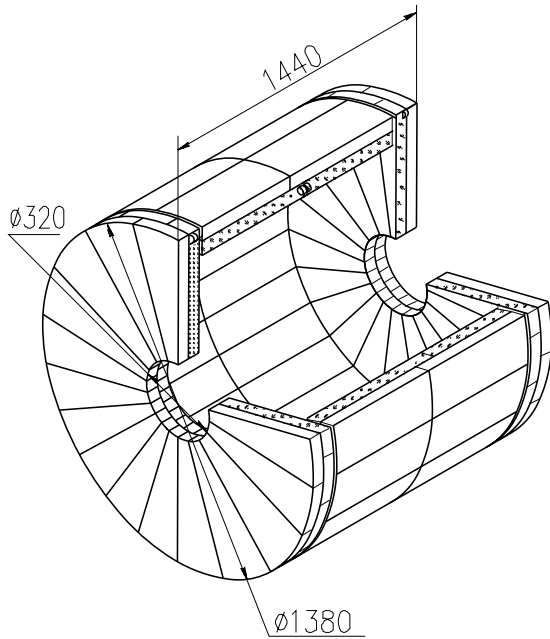
Pros:

- Large light collection area
- Small PMT (up to 10x smaller p.c. area in comparison with direct LC)
- Low cost

Cons:

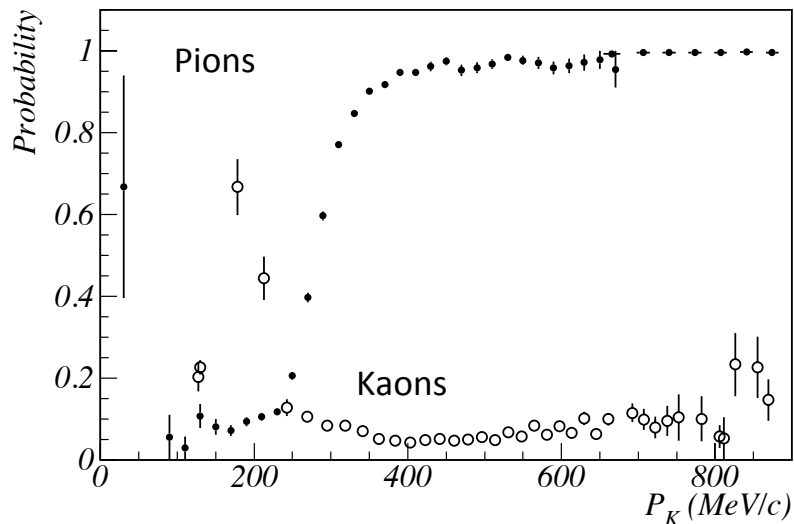
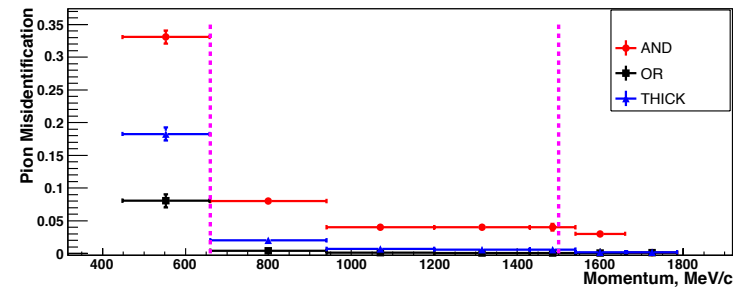
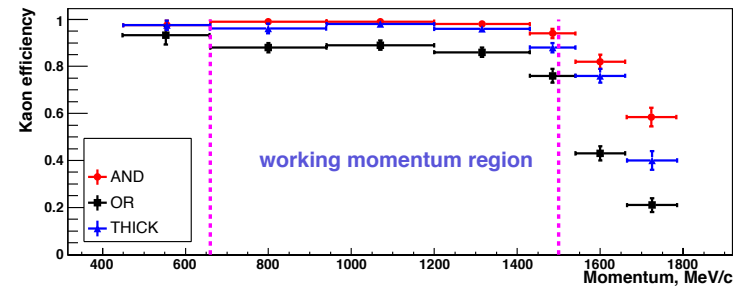
- Particle acceptance loss due to WLS

# Aerogel Cherenkov ASHIPH counters



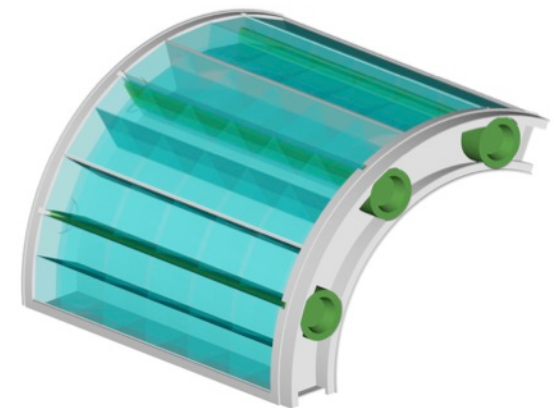
## **KEDR:**

- 160 counters (2 layers)
- $n=1.05$  (1000l)
- WLS (BBQ)
- MCP PMT  $\phi_{PC}=18$  mm
- $0.97 \times 4\pi$
- $24\% X_0$



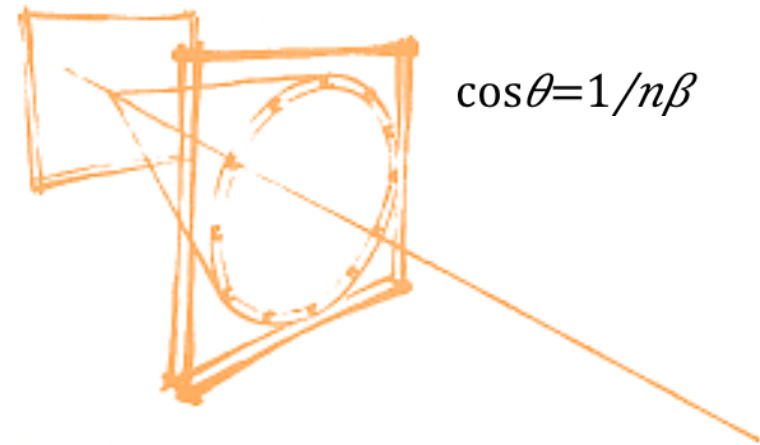
## **SND:**

- 9 counters (1 layer)
- $n=1.13$  (9l)
- WLS (BBQ)
- Thickness  $\sim 30$  mm
- MCP PMT  $\phi_{PC}=18$  mm
- $0.6 \times 4\pi$



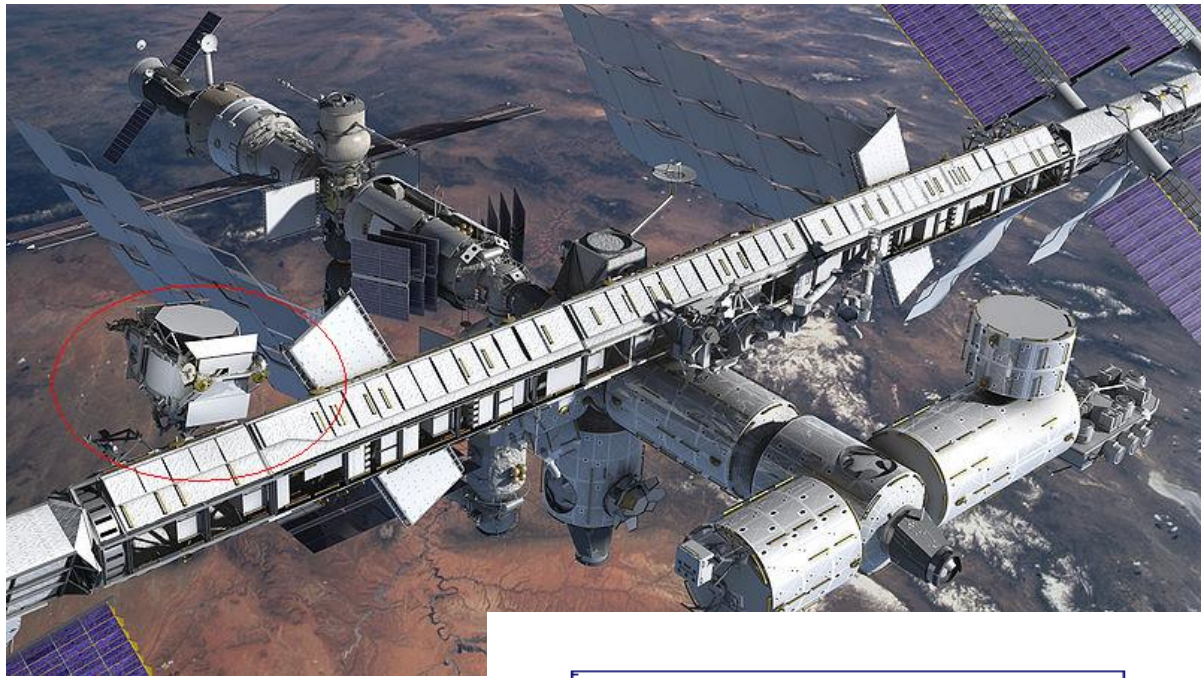
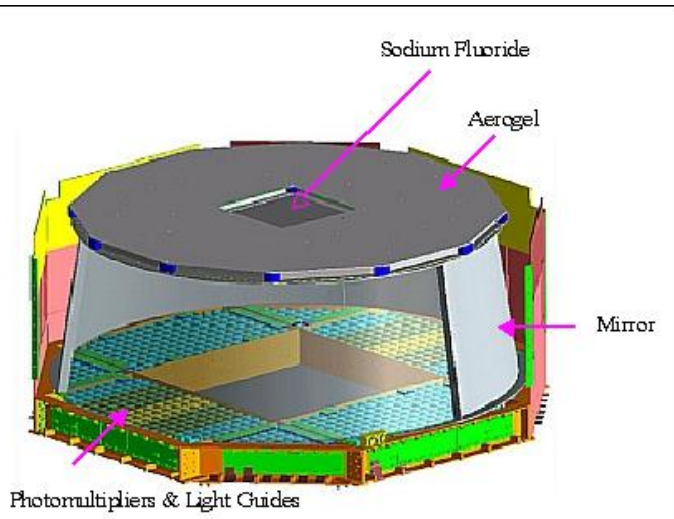
# Ring Imaging Cherenkov detectors with aerogel radiators

- If the Cherenkov radiation angle is measured, the precision in the determination (identification) of particle masses will be higher than in threshold counters.
- In the 1980s and 1990s, a whole series of RICH detectors were constructed:
  - CRID, SLD detector, SLAC(C6F14  $n=1.277$ , C5F12/N2  $n=1.0017$ )
  - RICH, Delphi detector, CERN, (C5F12|C6F14, C4F10)
  - RICH, CLEOIII detector, Cornell, (LiF,  $n=1.50$ )
  - DIRC, детектор BaBar, SLAC, США (SiO2,  $n=1.47$ )
- Main problem – they do not provide pion-kaon identification in the range of momenta 4–10 GeV/c
- **Material with  $n=1.03-1.05$  is needed.  
Aerogel!**



- A.Roberts, Nucl. Instrum. and Methods 9(1960)55
- J.Seguilot and T.Ypsilantis, Nucl. Instrum. and Methods 142(1977)377

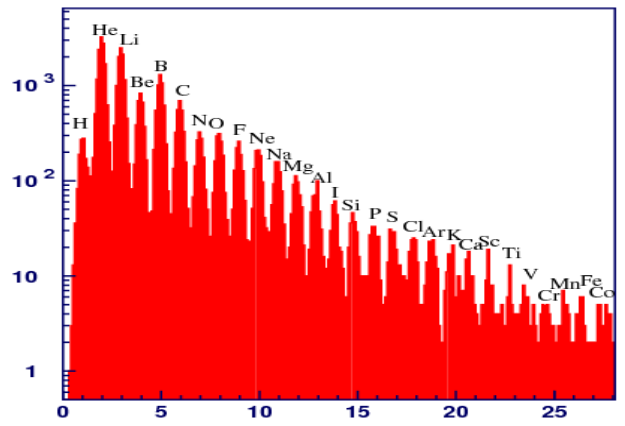
# Aerogel RICH of AMS-02 at ISS



- Antimatter search
- Dark matter
- Cosmic rays
- Strangelets search
- ...

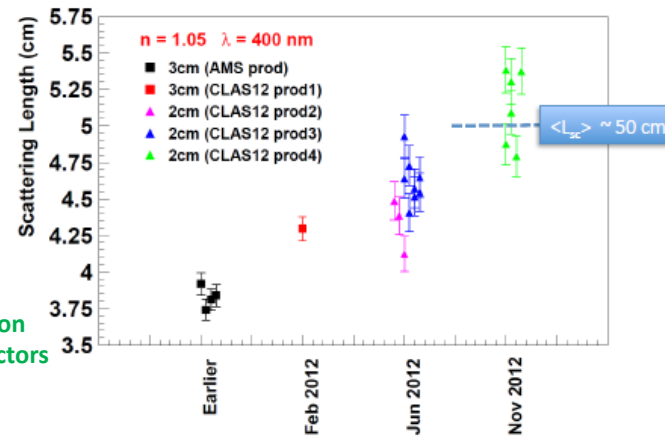
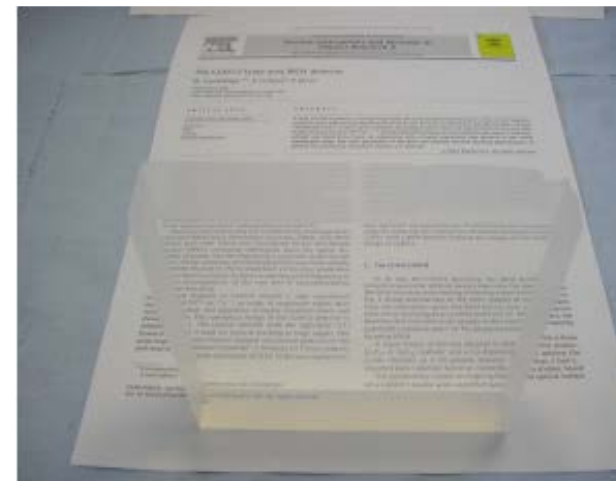
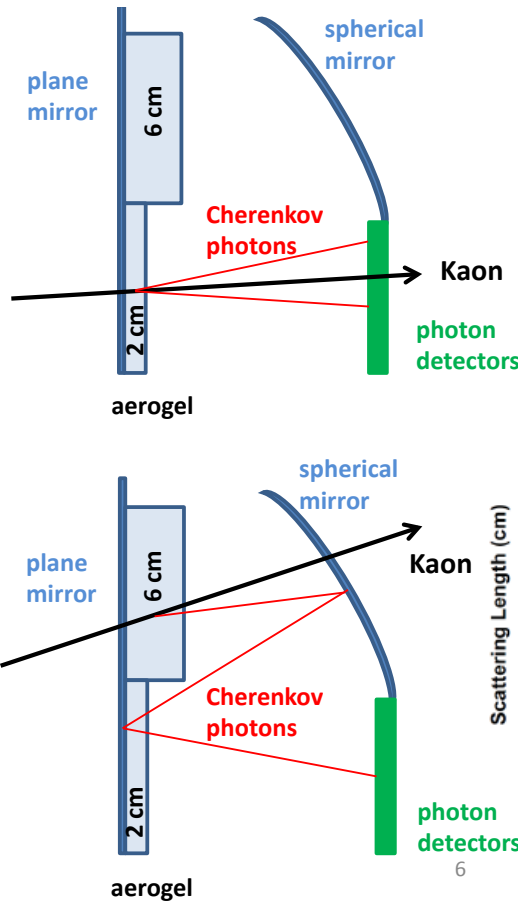
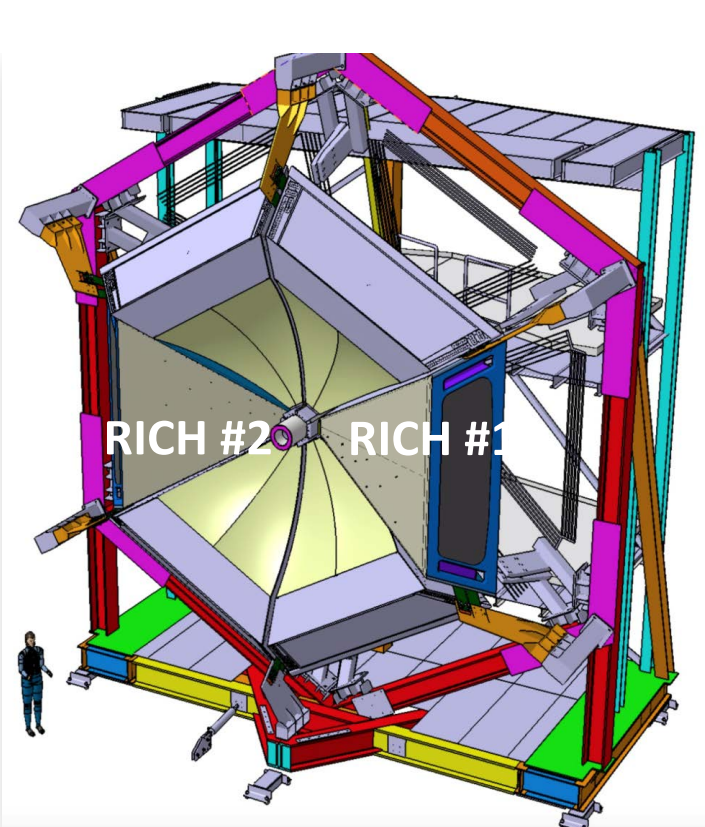
$M=8.5T$ ,  $V=54m^3$ ,  $S=1m^2$ ,  $B=1.26kGs$   
 TOF, TRD, RICH, Si Tracker, eCal

Measurement of  $Z$  of the nucleon,  $N_{pe} \sim Z^2$   
 BIC/BINP production,  $n=1.05$   
 It has working at ISS since 2011





# RICH detector for the CLAS12



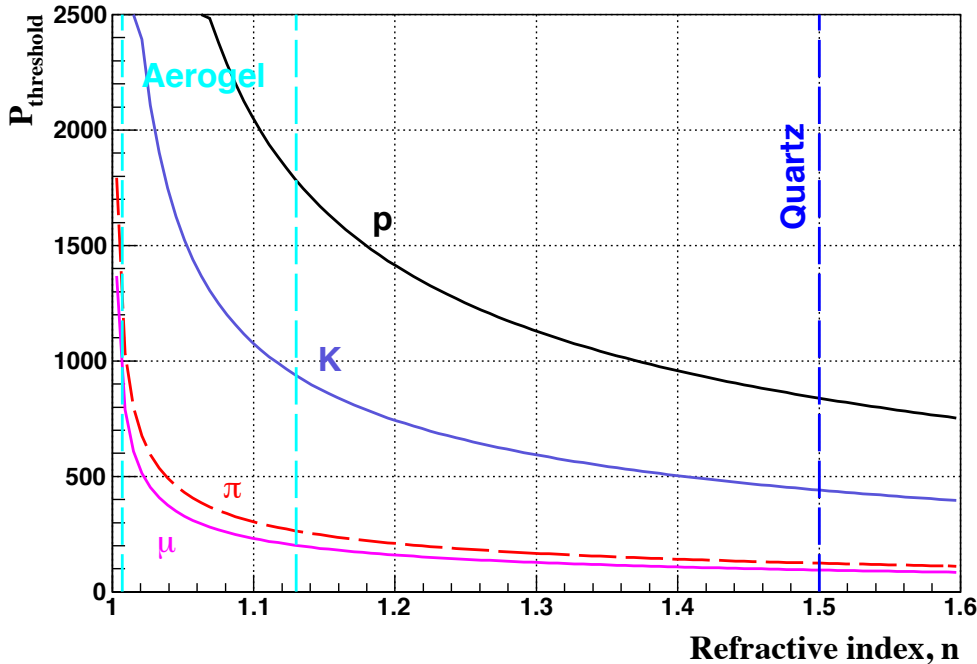
- **K/π** and **K/p** separation at  **$>4\sigma$**  level in **few GeV/c** region;
- The RICH will replace the existing LTC detectors;
- Installation in CLAS12 by **September 2017**;

*RICH 2016, September 9th 2016, Bled: talks by M.Marazita, M.Contralbrigo, E.Kravchenko*

# Focusing Aerogel RICH for PID system

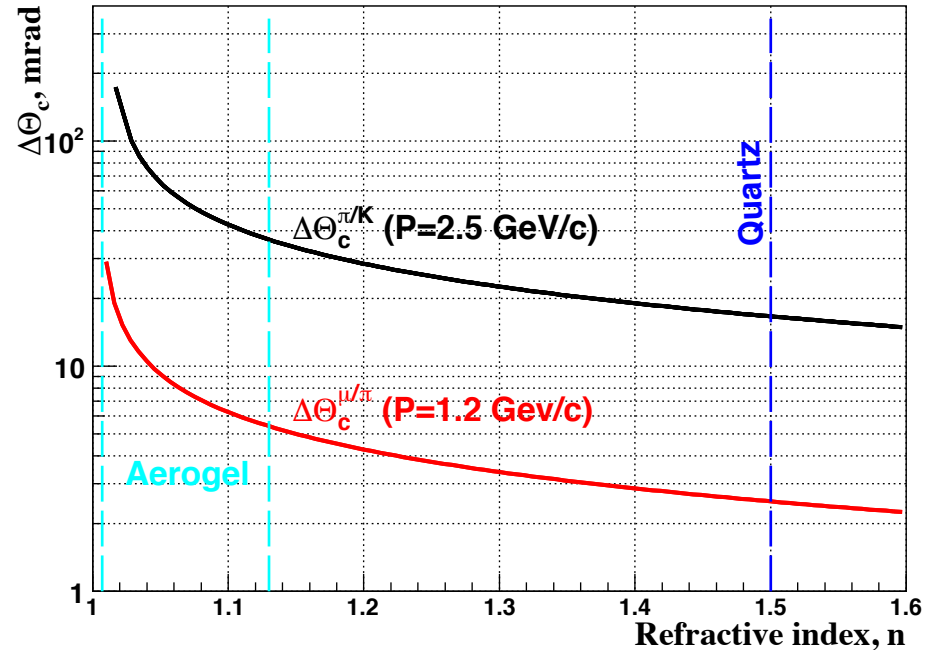
## (Motivation)

Dependence of Cherenkov threshold momentum on refractive index



Aerogel threshold counters could be used for reliable  $K/\pi$  separation.

Dependence of  $\Delta\theta_c$  on refractive index

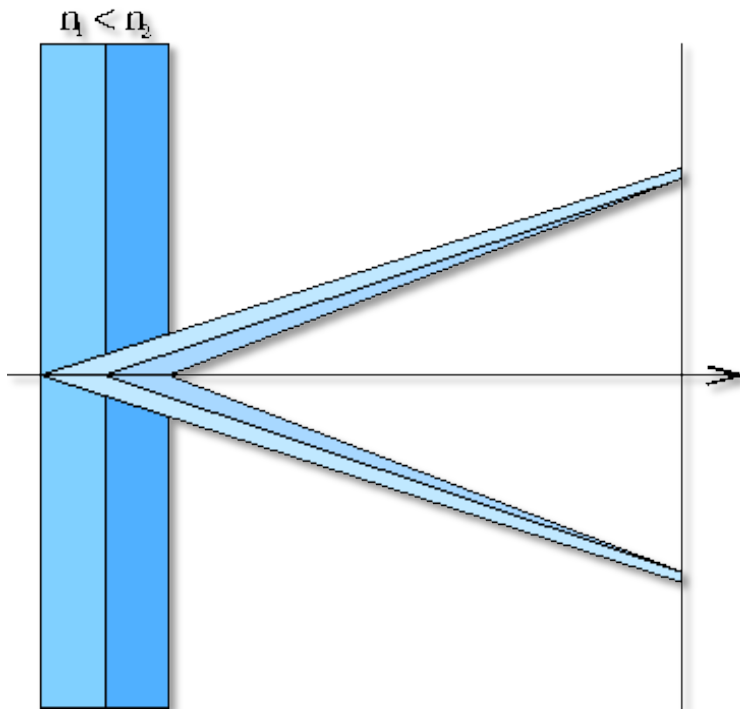


For reliable  $\mu/\pi$  separation we need aerogel RICH with  $\sigma_{\text{track}}(\theta_c) < 2.5$  mrad

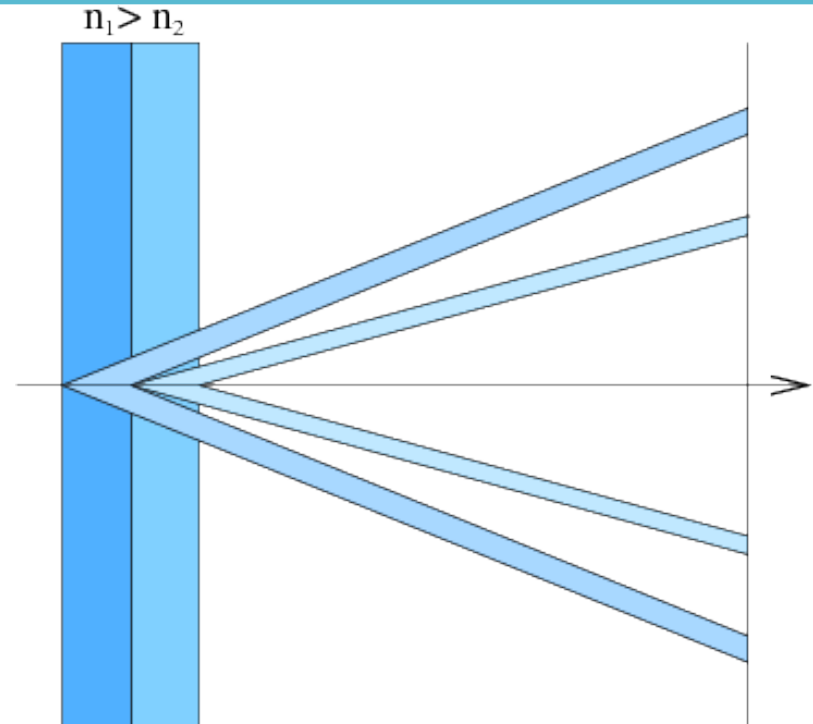
# Focusing Aerogel RICH – FARICH (Concept)

Focusing aerogel improves proximity focusing design by reducing the contribution of radiator thickness into the Cherenkov angle resolution

## Single ring option



## Multi-ring option



T.Iijima et al., NIM A548 (2005) 383

A.Yu.Barnyakov et al., NIM A553 (2005) 70

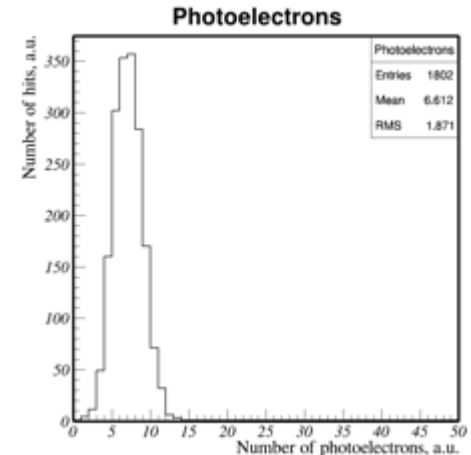
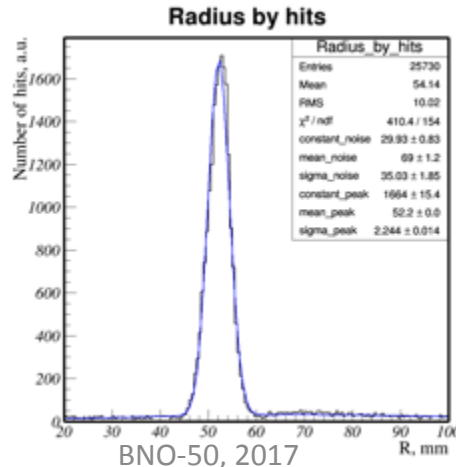
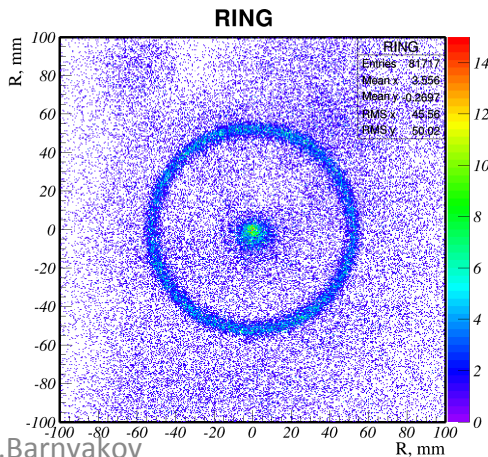
# Single ring option: two approaches

## Two blocks

- Aerogel RICH for Belle-II:
  - $n_1=1.045$ ,  $n_2=1.055$
  - Thickness – 20 + 20 mm
  - Distance – 200 mm
- HAPD with 5x5 mm pixel
- $\sigma_\theta=15.8$  mrad and  $N_{pe}=8.6$   
 $\sigma_\theta(\text{track}) = \sigma_\theta / \sqrt{N_{pe}} \approx 5.4$  mrad  
*S.Nishida et al., NIM A 766 (2014) 28*

## Two layer block

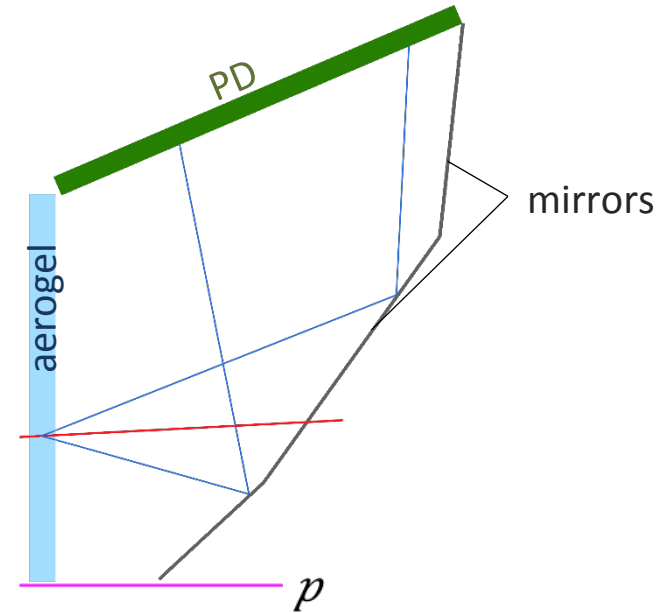
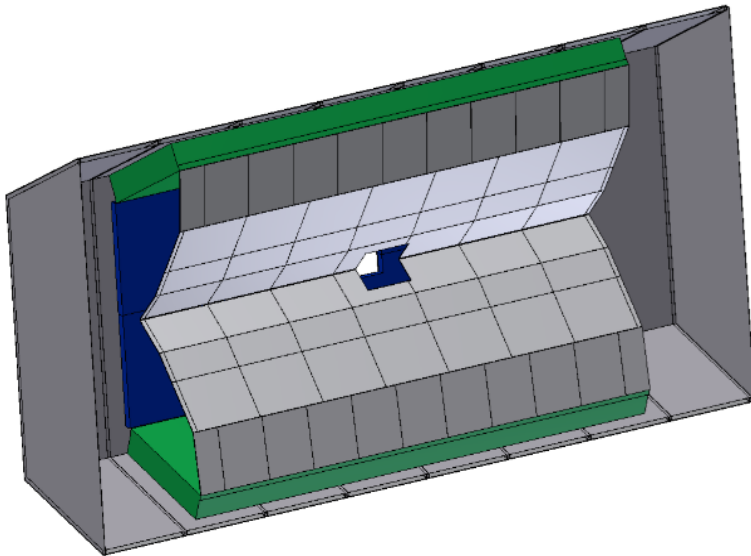
- Aerogel from BINP&BIC:
  - $n_1=1.045$ ,  $n_2=1.053$
  - Thickness – 15 + 15 mm
  - Distance – 200 mm
- Philips DPC3200 – 4x4 mm pixel
- $\sigma_\theta=11.2$  mrad and  $N_{pe}=6.6$   
 $\sigma_\theta(\text{track}) = \sigma_\theta / \sqrt{N_{pe}} \approx 4.4$  mrad  
*Preliminary results of BINP testbeam 2016*



# Forward Spectrometer RICH for PANDA

(BINP, BIC)

## PANDA Forward RICH



Particle ID:  $\pi/K/p$  up to 10 GeV/c  
3m<sup>2</sup> detector area (MaPMTs or SiPMs)

### Radiator

- Focusing 2- or 3-layer aerogel
- 40 mm thick

### Photon Detector

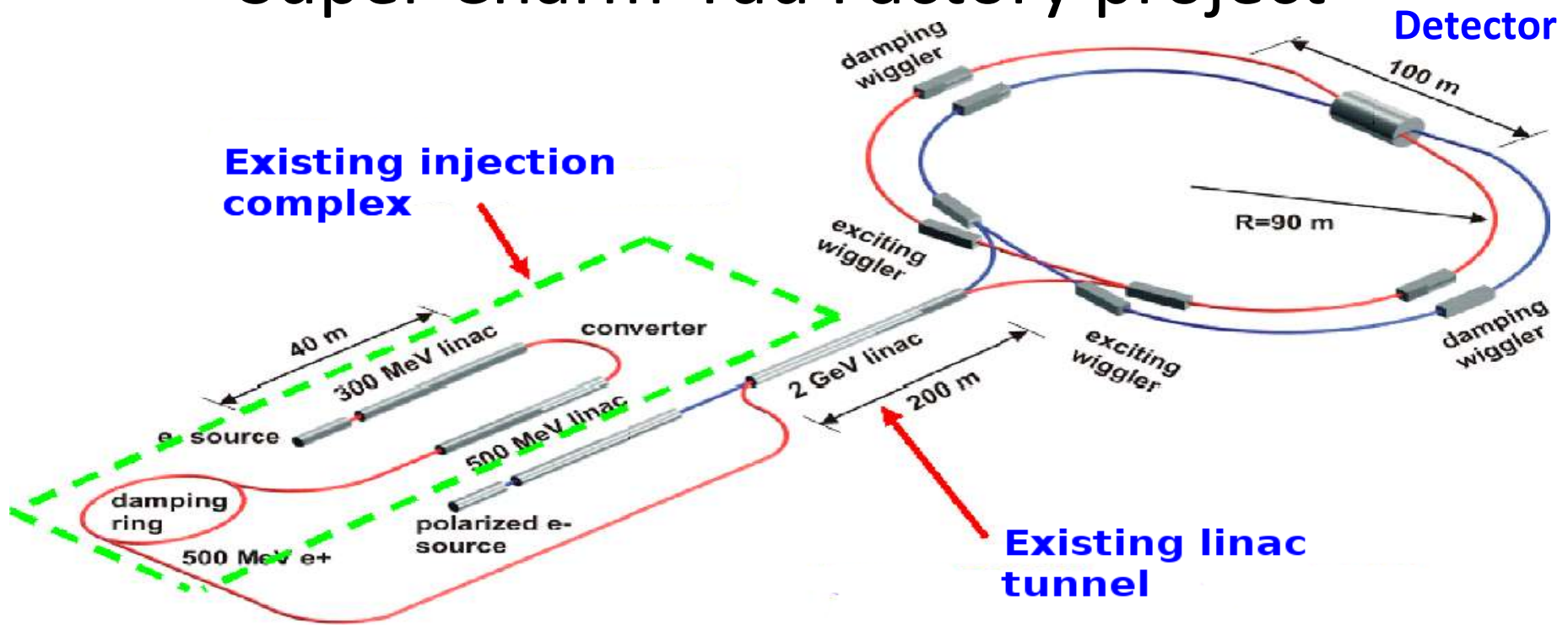
Hamamatsu H12700 MaPMT

- flat panel,
- 8x8 anode pixels of 6mm size
- 87% active area ratio
- Bialkali photocathode
- Gain:  $1.5 \cdot 10^6$

### Mirrors

- Flat segments
- Float glass substrate 2 mm thick
- Al+SiO<sub>2</sub> coating, R $\geq$ 90%

# Super Charm-Tau Factory project



## Factory outline:

- Double ring Symmetric e<sup>+</sup>e<sup>-</sup> collider with Crab Waist scheme
- $E_{\text{c.m.}} = 2 \div 5 \text{ GeV}$
- $L = 10^{35} \text{ cm}^{-2}\text{s}^{-1}$   
(100 times more than existing c- $\tau$  factories)
- Longitudinal polarization of e<sup>-</sup> - beams in IP

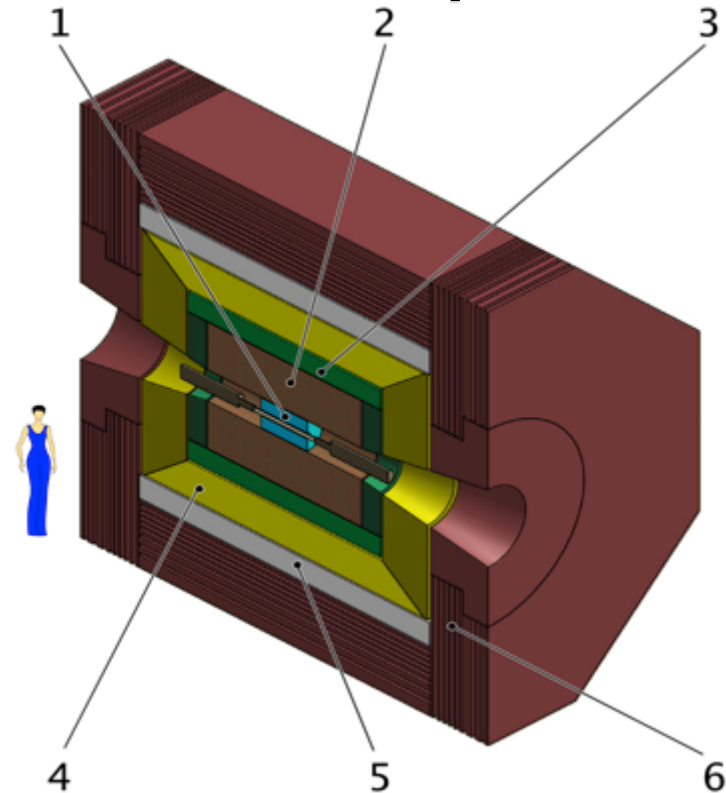
# Detector for Super CT-factory

## Physical program:

- Rare decays of D mesons,  $\tau$  lepton;
- $D^0\bar{D}^0$  oscillations;
- Searches for lepton-flavor-violating decays of  $\tau$  (for instance  $\tau \rightarrow \mu\gamma$ );
- ...

## Detector requirements

- An excellent momentum resolution for charged particles and a good energy resolution for photons;
- $K/\pi$  separation higher than  $3\sigma$ ;  
 $\mu/\pi$  separation up to  $1.5 \text{ GeV}/c$ ;
- DAQ system, which is able to read events at a rate of  $300\div 400\text{kHz}$  with  $30\text{kB}$  event length;
- ...

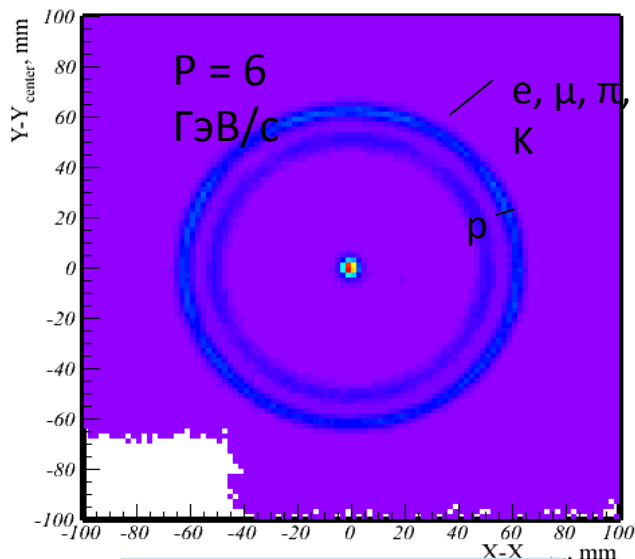


- 1 – Vertex Detector
- 2 – Drift Chamber
- 3 – PID => FARICH
- 4 – EMC
- 5 – Superconducting Solenoid
- 6 – IFR

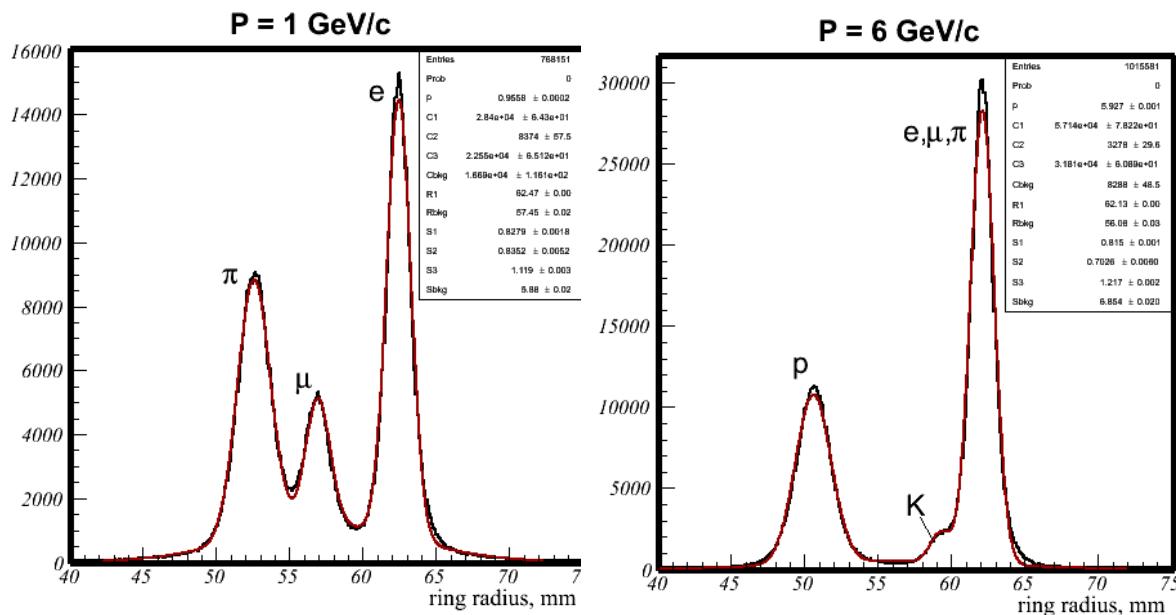
See CTF CDR ([https://ctd.inp.nsk.su/docs/ScTau\\_CDR\\_en/CDR\\_en\\_ScTau.pdf](https://ctd.inp.nsk.su/docs/ScTau_CDR_en/CDR_en_ScTau.pdf))

# Beam test results at CERN PS T10, June 2012

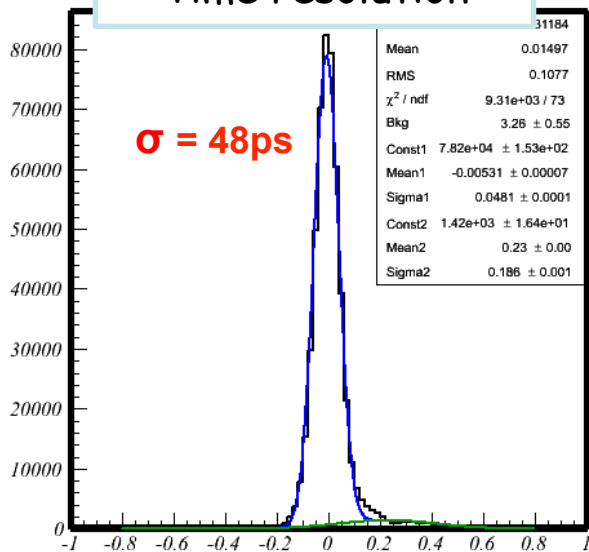
## Hit positions



## Ring radius distributions



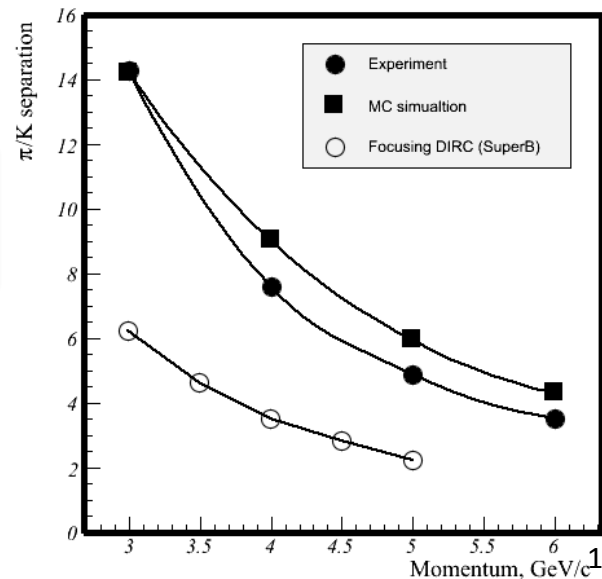
## Time resolution



$$S(\pi/K) = \frac{R_{\pi} - R_K}{\sigma_{\pi}}$$

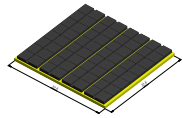
$\pi / K : 7.6\sigma @ 4 \text{ GeV/c}$   
 $\mu / \pi : 5.3\sigma @ 1 \text{ GeV/c}$

A.Yu. Barnyakov, et al.,  
 NIM A 732 (2013) 352

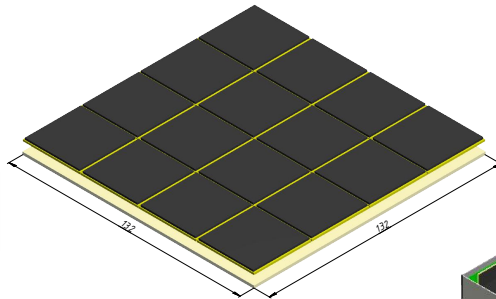




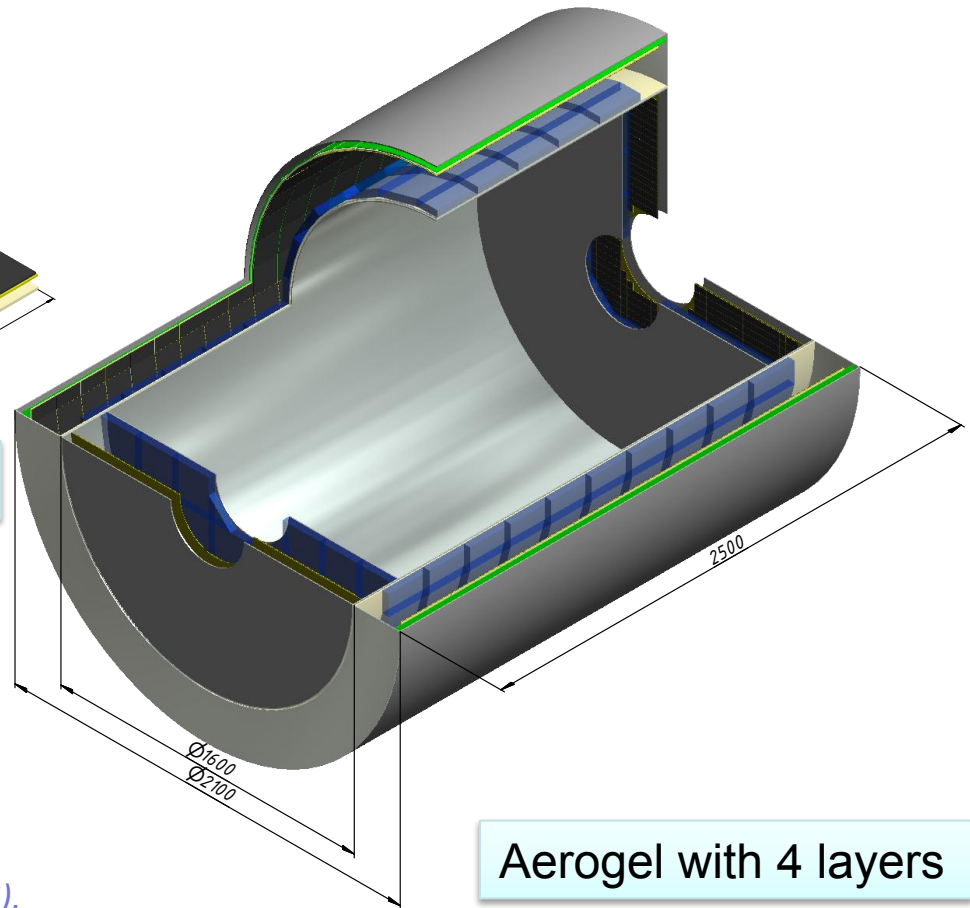
# FARICH system



Tile with 64 pixels



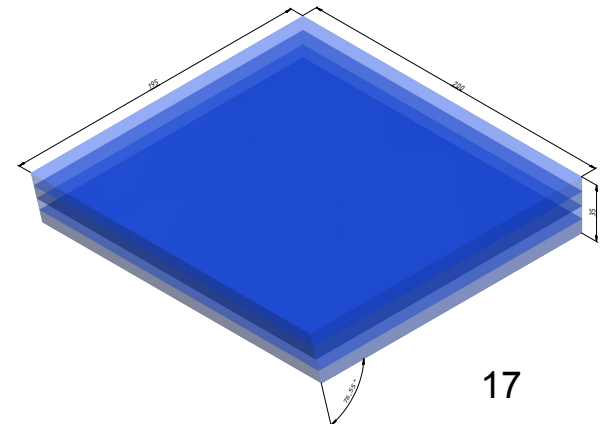
Module with 16 tiles



Aerogel with 4 layers

## Main parameters

- Focusing aerogel radiator,  $n_{\max}=1.07$ , 4 layers
- Photon detector:  $\sim 3 \times 3 \text{ mm}^2$ , pitch 4 mm  
*DPC (Philips), MPPC (Hamamatsu), NUV-HD (FBK-IRIS), Array-C (SensL)*
- Area of the photon detector:  $20 \text{ m}^2$
- Area of the radiator:  $14 \text{ m}^2$
- $\sim 10^6$  channels



# Summary

- Development of aerogel Cherenkov counters for HEP experiments have been carrying out in Novosibirsk since 1986 by BINP and BIC in close cooperation.
- The ASHIPH method for threshold Cherenkov counters was developed. Good  $\pi/K$ -separation was achieved in ASHIPH system of the KEDR detector and the SND detector.
- Aerogel radiators for RICH detector of the AMS02 experiment were produced in Novosibirsk. Experimental number of photoelectrons and Cherenkov angle resolution are in good agreement with MC simulation. The AMS02 experiment has been operating since 2010 at ISS.
- Production of aerogel radiators with 200x200 transvers dimensions and high transparency for RICH of CLAS12 project is organized in Novosibirsk.
- Two projects of FARICH detectors for PANDA experiment and SCTF experiment are under development now.

# Addendum

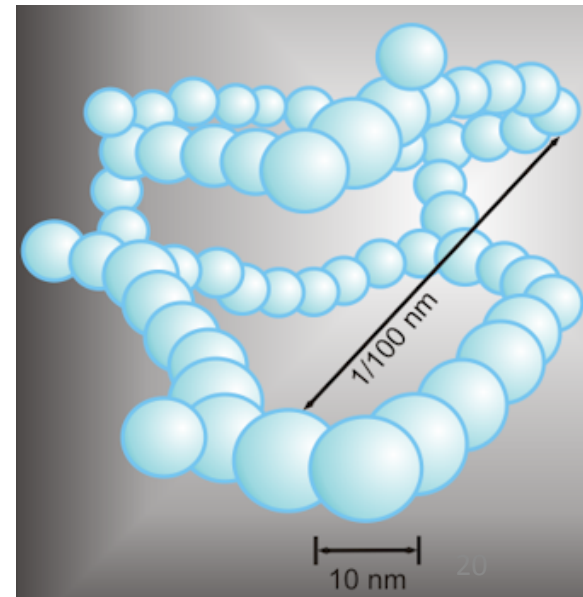
# What it is -- Aerogel?(1)



- Aerogel was first synthesized by Samuel Stephens Kistler in 1931
- S.S.Kistler, "*Coherent Expanded Aerogels and Jellies*", Nature, 1931,vol. 127, p. 741

**Aerogel** – is a porous material with pore dimension less than visible light wavelength.

It is a classical **nanomaterial**.The most widespread are silicon dioxide aerogel, although aerogels based on metal oxides, carbon, gelatin and others.



# What it is -- Aerogel?(2)

- Production method:
- Synthesis of the alcogel:
  - $\text{Si(OR)}_4 + 2\text{H}_2\text{O} \Rightarrow \text{SiO}_2 + 4\text{HOR}$
  - alkoxide    water    silica    alcohol
- Supercritical drying in the autoclave to remove alcohol  $P_{\text{max}}=100 \text{ atm}$ ,  $T_{\text{max}}=260^\circ\text{C}$ 
  - methanol --  $P_{\text{cr}}=81 \text{ atm}$ ,  $T_{\text{cr}}=230^\circ\text{C}$
  - isopropanol --  $P_{\text{cr}}=53 \text{ atm}$ ,  $T_{\text{cr}}=235^\circ\text{C}$
  - carbon dioxide --  $P_{\text{cr}}=73 \text{ atm}$ ,  $T_{\text{cr}}=31^\circ\text{C}$
- Aerogel parameters:
  - Density 0.003 до 1.0 g/cm<sup>3</sup> (*fused silica*  $\rho=2.2 \text{ g/cm}^3$  )
  - Refractive index
    - $n \approx 1 + 0.2 \cdot \rho[\text{g/cm}^3] \Rightarrow$
    - $(n = 1.0006 \div 1.2)$
  - Porosity 99.8%
  - Inner surface 800 m<sup>2</sup>/g

# Era of high transparency aerogel

L.W.Hrubesh, T.M. Tillotson, J.F. Poco “*Characterization of ultralow-density silica aerogels made from a condensed silica precursor*”, MRS Proc. 180(1990)315

- One-step technology
- Direct alcogel synthesis
- $\text{Si(OR)}_4 + 2\text{H}_2\text{O} \Rightarrow \text{SiO}_2 + 4\text{HOR}$
- alkoxide    water    silica    alcohol
- $L_{\text{sc}}(400) \sim 20 \text{ mm}$
- Two-step technology
- A mixture of oligomers preparation
- $\text{Si}_k\text{O}_l(\text{OR})_m(\text{OH})_n \Rightarrow \text{SiO}_2 + \text{alcohol}$
- $L_{\text{sc}}(400) > 35 \text{ mm}$

**Two-step technology was implemented at BIC in 1992**

# Aerogel threshold counters with wavelength shifters(1)

- At  $\lambda=400$  nm

- $L_{sc} \sim 40$  mm,  $L_{abs} \sim 4-5$  m

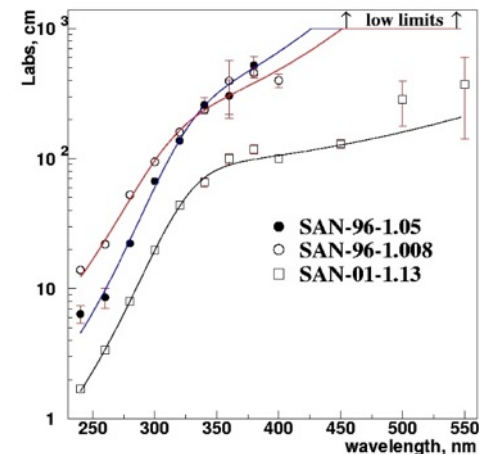
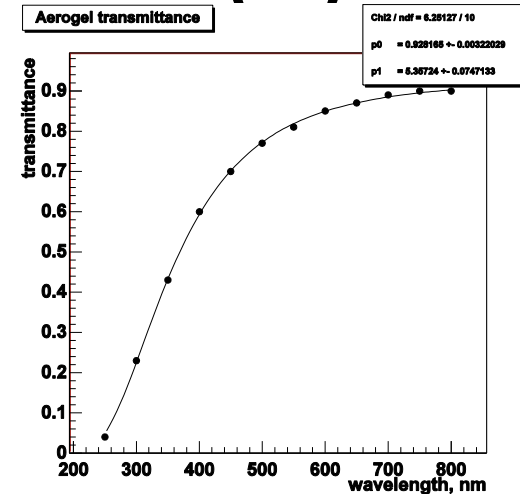
- At  $\lambda=300$  nm

- $L_{sc} \sim 12$  mm,  $L_{abs} \sim 0.5-1$  m

–But!

- $dN/d\lambda \sim 1/\lambda^2$
- At  $\sim 300$  nm Number of Cherenkov photons is 3 times larger than at  $\sim 400$  nm

–The idea is to absorb Cherenkov photons at short wavelengths and re-emit at large where aerogel transparency is better.

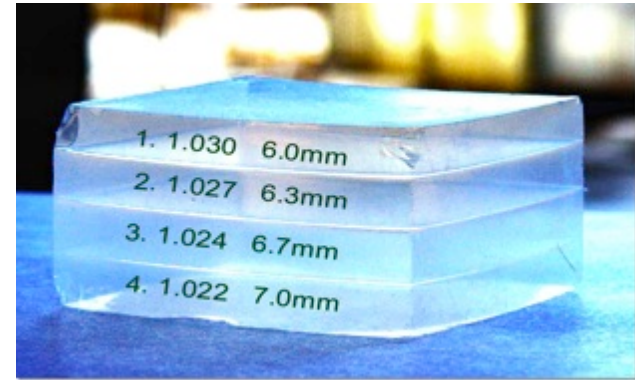


# Status of aerogel production in Novosibirsk

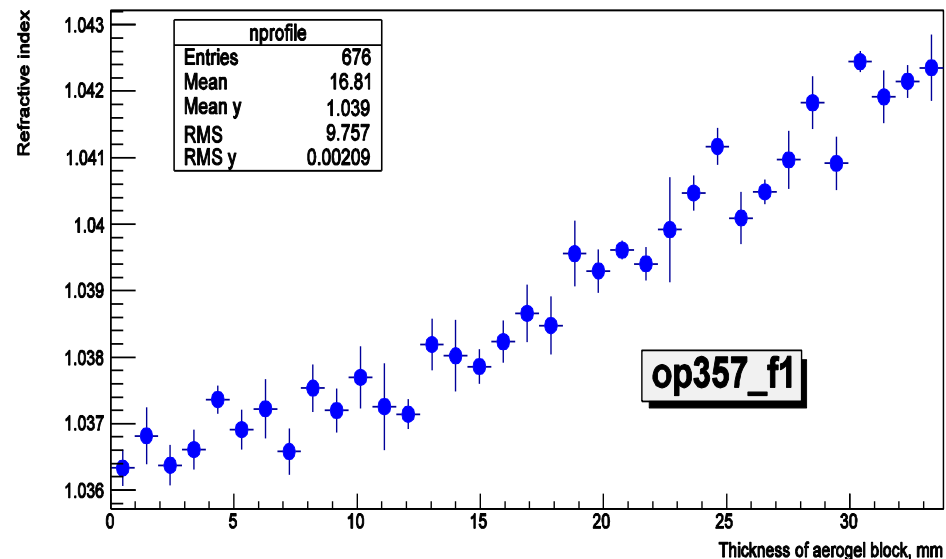
- Aerogel production in Novosibirsk was started in 1986.
- $n=1.006\div 1.13$ ;  $L_{sc}(400nm) \geq 43$  mm.
- In 2004 first 4-layer tile was produced
- 2,3,4-layer blocks with  $n_{max}=1.05$  100x100x30 mm were produced in recent years.
- Tiles with  $n=1.05$  and 200x200x30 mm are produced for J-Lab CLAS12 experiment. Total amount is 6 m<sup>2</sup>.
- In 2012 development of aerogel production with continuous designed profile of density gradient was started.

## The aims:

- Regular production of 3,4-layer tiles with  $n_{max}=1.07$  and 200x200x35 mm.
- Development of “gradient” aerogel production.



*A. Yu. Barnyakov et al., NIM A553 (2005) 70*

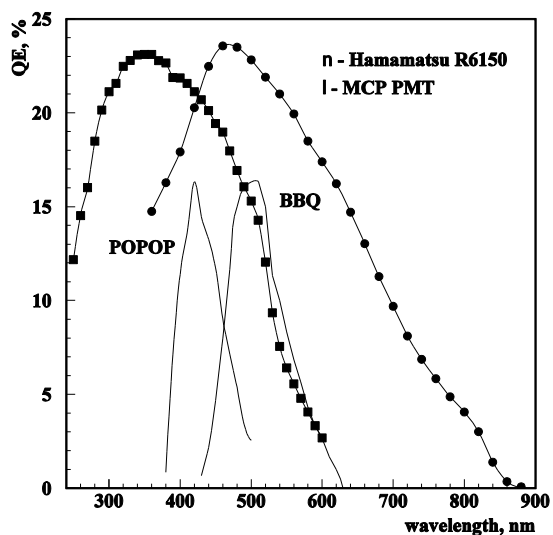
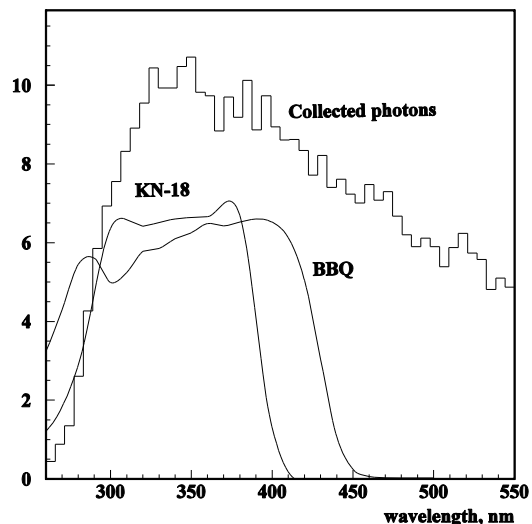


*A. Yu. Barnyakov et al., NIM A766 (2014) 88*

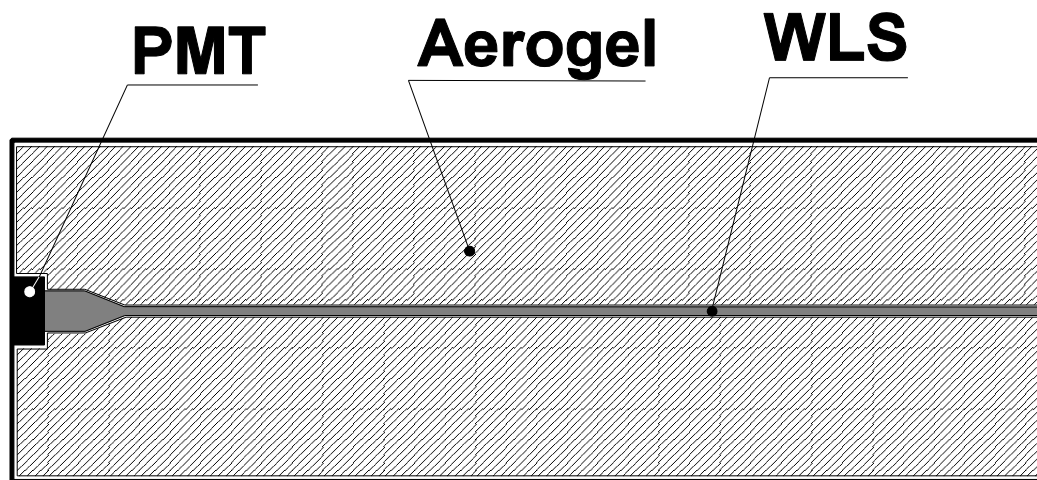
*A. Yu. Barnyakov et al., NIM A766 (2014) 235*



# ASHIPH detectors



## Aerogel Shifter and Photomultiplier

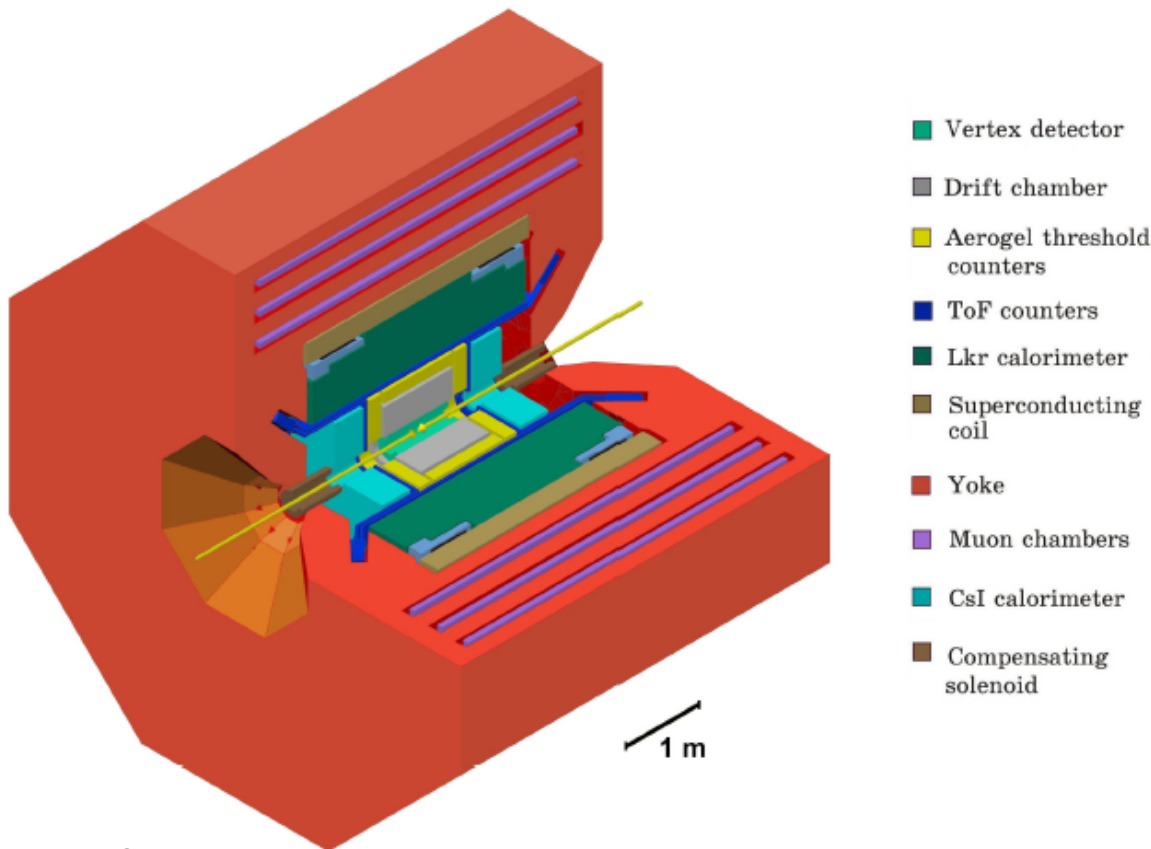


PMMA light guide doped with BBQ dye is used as wavelength shifter

Suggested at BINP. A.Onuchin et.al. NIM A315(1992)517



# KEDR experiment at VEPP-4M



- Vertex detector
- Drift chamber
- Aerogel threshold counters
- ToF counters
- Lkr calorimeter
- Superconducting coil
- Yoke
- Muon chambers
- CsI calorimeter
- Compensating solenoid

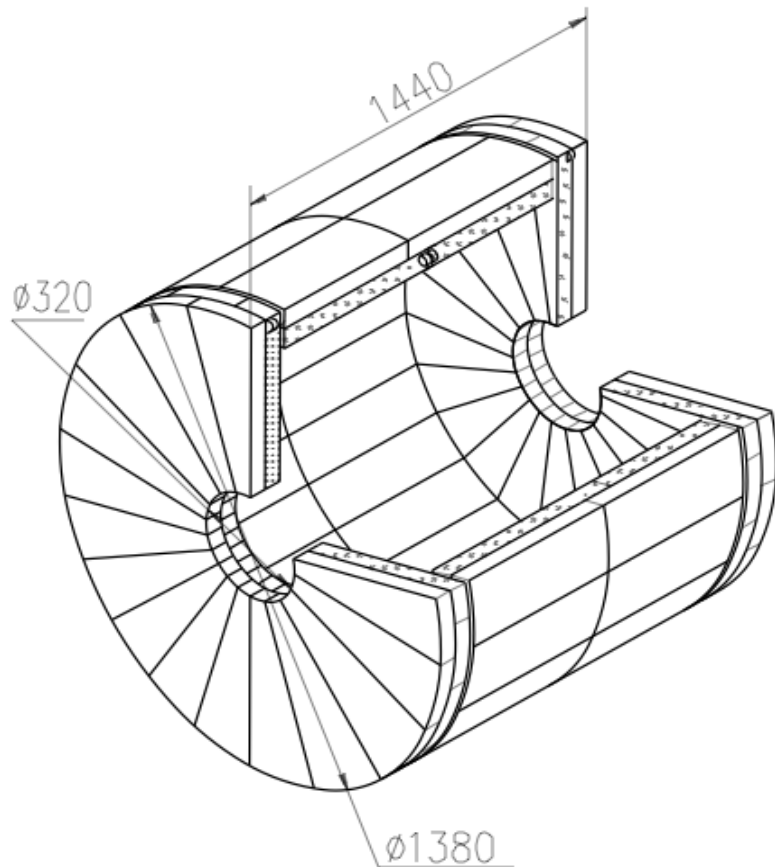
## VEPP-4M

- Symmetric  $e^+e^-$  collider
- $E_{c.m.} = 2-10$  GeV
- $L = (1 \div 80) \times 10^{30} \text{ cm}^{-2}\text{s}^{-1}$
- Precise energy calibration:  
RD:  $(5 \div 15) \times 10^{-6}$   
CBS:  $3 \times 10^{-5}$

## Physics program

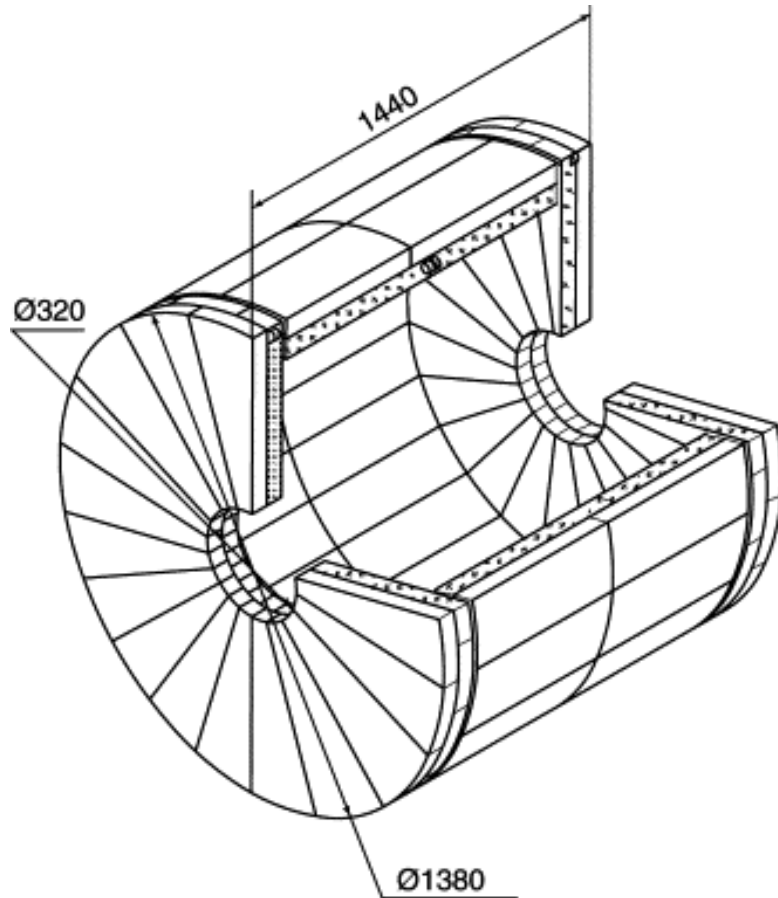
- Precise particle mass measurements:  $J/\psi$ ,  $\psi(2S)$ ,  $\psi(3770)$ ,  $\tau$  lepton,  $D$  mesons,  $Y$  mesons
- Measurements of  $\psi$  and  $Y$  mesons lepton width
- R measurement in 2-10 GeV c.m. energy range
- $\gamma\gamma \rightarrow \text{hadrons}$  and other  $2\gamma$  processes
- Branching fractions measurements in charm and bottom quark systems (above  $10^{-4}$ )

# KEDR ASHIPH system(1)



- 160 counters in 2 layers
- Solid angle 96% of  $4\pi$
- $n=1.05$ ,  $V_2=1000$  l, high transparency SAN-96 aerogel
- $\pi/K$ - separation in the momentum range  $0.6 \div 1.5$  GeV/c
- 160 MCP PMTs, photocathode diameter  $\varnothing 18$ mm, able to work in the magnetic field up to 2 T
- Fully installed in the detector in 2013. Now in operation.

# KEDR ASHIPH system



- $n=1.05$
- $\pi/K$  separation:  
0.6 – 1.5 GeV/c
- 160 counters in 2 layers
- 160 MCP PMTs
- works in magnetic field up to 1.5T
- 1000 liters of aerogel
- 24%  $X_0$  for 2 layers

## Beam test in Dubna (2000)

A.Yu.Barnyakov et al., NIM A494 (2002) 424

$N_{pe} \approx 10$  at saturation

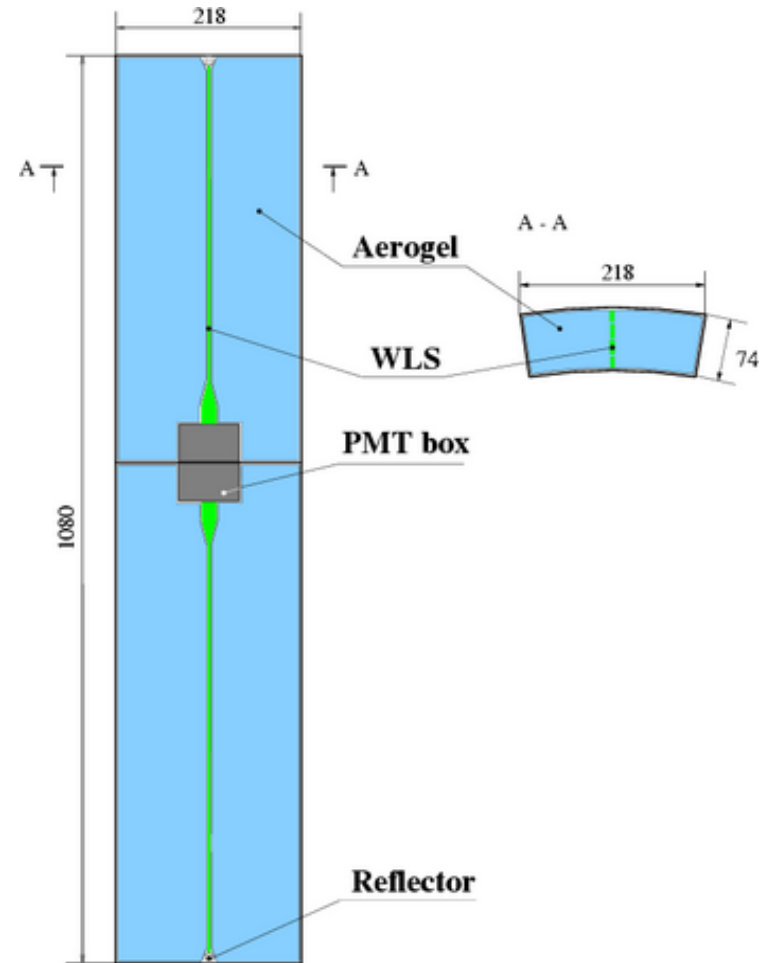
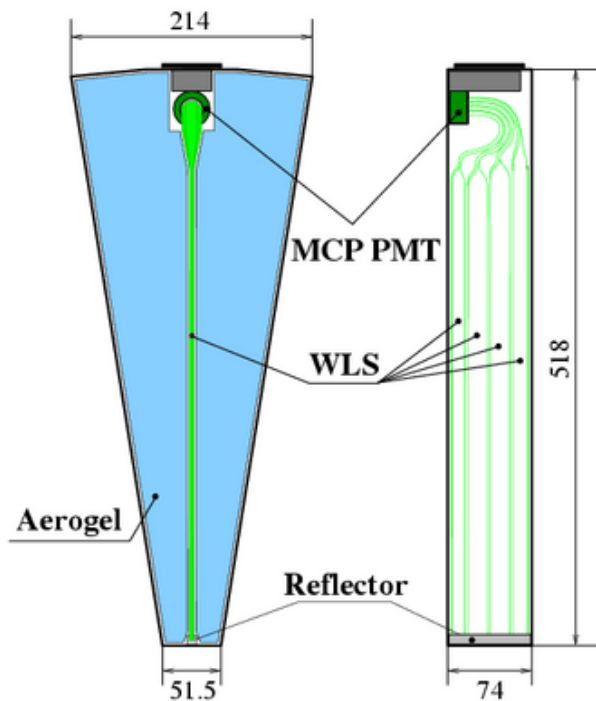
P=0.86 GeV/c	P=1.2 GeV/c
pion rejection 900:1 kaon efficiency 0.94 $\pi/K$ separation <b>4.7 <math>\sigma</math></b>	pion rejection 1300:1 kaon efficiency 0.90 $\pi/K$ separation <b>4.5 <math>\sigma</math></b>



# KEDR ASHIPH counters

## Barrel counter

## End-cap counter



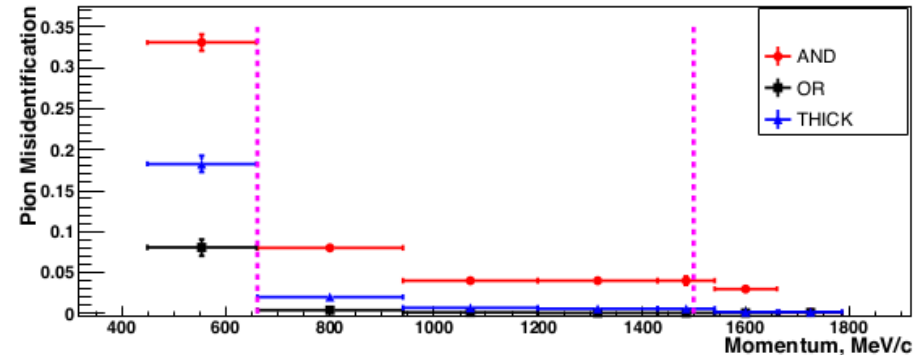
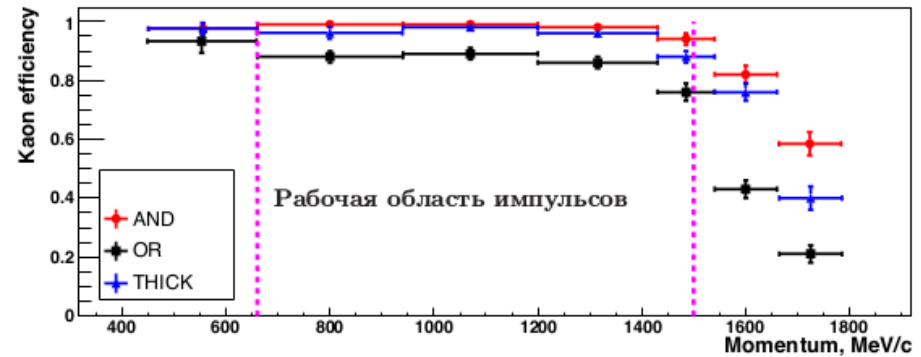


A.Yu.Barnyakov

BNO-50, 2017

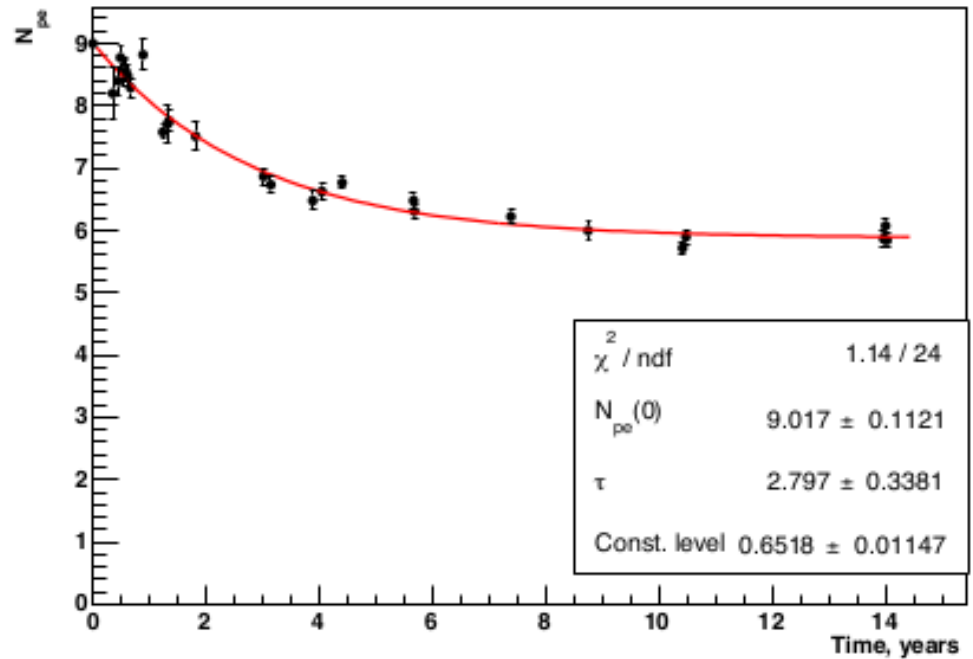
# KEDR ASHIPH system(2)

- $N_{pe} = 6.4 \pm 0.2$  – layer 1
- $N_{pe} = 5.0 \pm 0.2$  – layer 2
- $N_{pe} = 10.9 \pm 0.2$  – sum of the signals in 2 layers (80%)
- $\pi/K$  separation at  $1.2\text{GeV}/c$  is  $4.3\sigma$



# ASHIPH long term stability

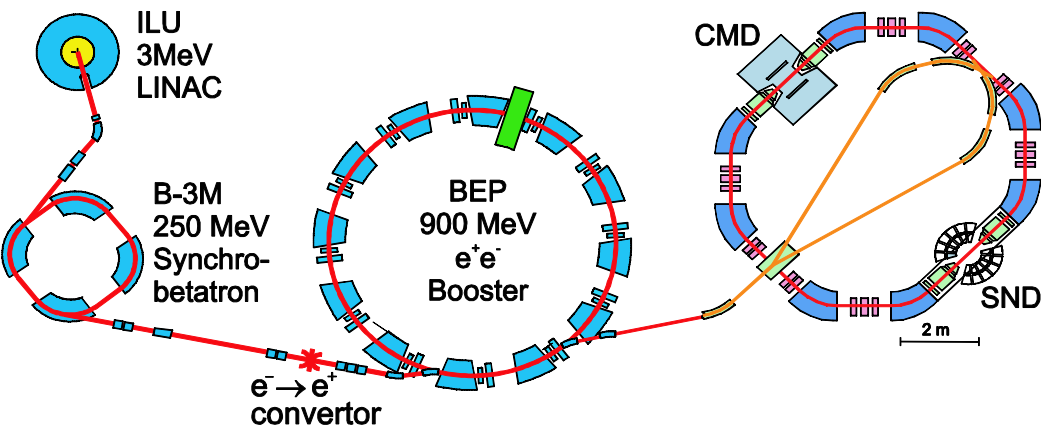
- A prototype of the endcap ASHIPH counter are under operation since 2000. From time to time it is tested in Cosmic Ray Telescope (CRT). Its signal degradation now has stabilized at the level of 60% from initial value.





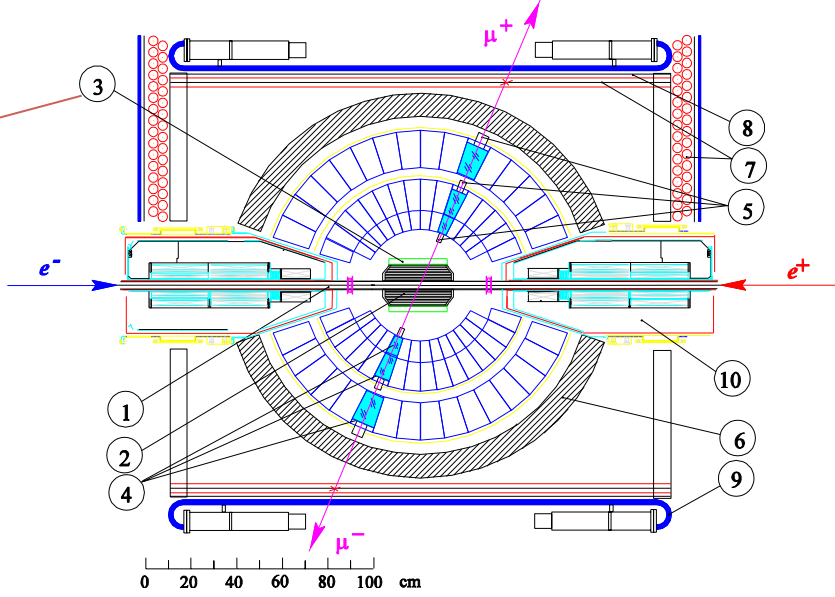


# SND at VEPP-2000



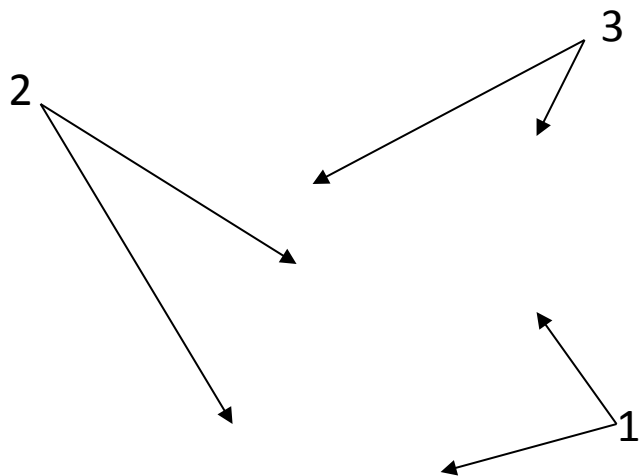
Symmetric  $e^+e^-$  collider with round beams  
 $2E_{max}=2000$  MeV  
 $L=10^{31} \text{cm}^{-2}\text{s}^{-1}$  at  $E=510$  MeV  
 $L=10^{32} \text{cm}^{-2}\text{s}^{-1}$  at  $E=1000$  MeV

1. VEPP-2000 beam pipe
2. Tracking system
3. Aerogel Cherenkov counters
4. NaI(Tl) crystals
5. Vacuum phototriodes
6. Fe absorber
7. } EMC
8. } Muon system
9. }
10. VEPP-2000 s.c. focusing solenoids





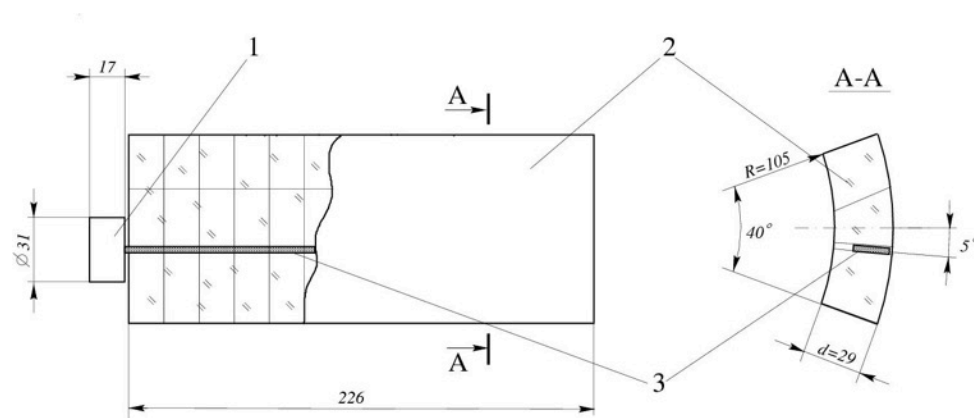
# Aerogel Cherenkov counters for SND



1 - PMT, 2 – aerogel, 3 - WLS

## Outline

- $\pi/K$  separation from 300 to 870 MeV/c
- Cylindrical shape:  $R=105\div 141$  mm
- Case material: 1mm of Al
- 3 segments of 3 counters in each
- Solid angle:  $\sim 60\%$  of  $4\pi$
- Thickness:  $0.09 X_0$



## Counter Design

- Scheme: ASHIPH
- WLS position: displaced by  $\sim 5^\circ$  from counter center
- Aerogel thickness:  $\sim 30$  mm

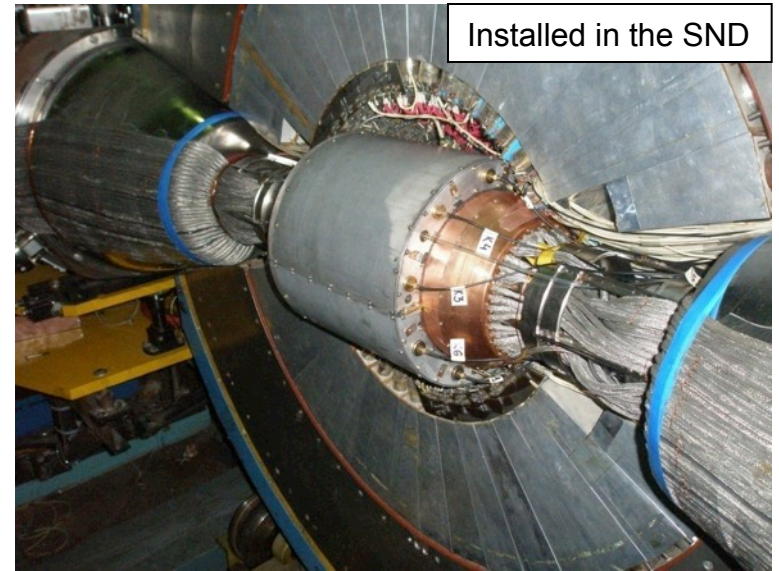
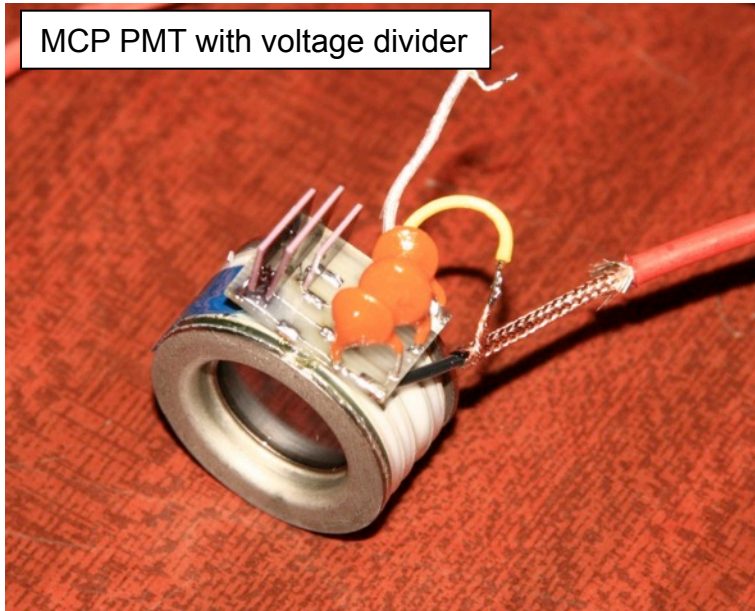
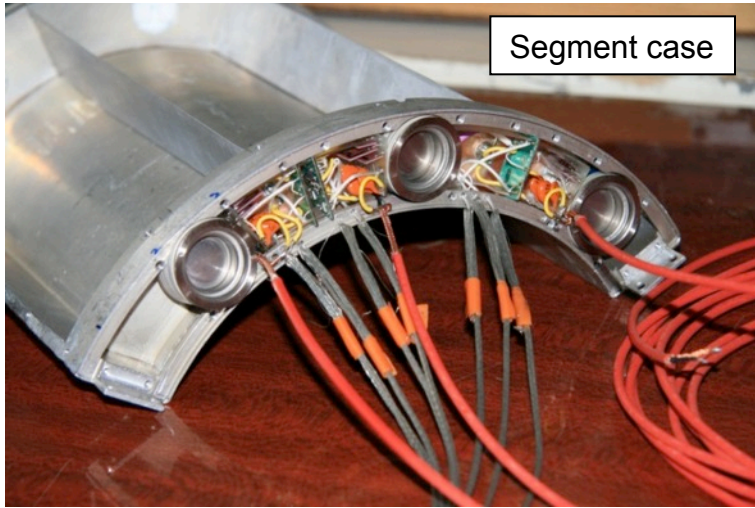
## Aerogel

### parameters

- Refraction index:  $n=1.13\pm 0.01$
- Density:  $\rho=0.65$  g/cm<sup>3</sup>
- $L_{sc}=19$  mm at  $\lambda=400$  nm
- $L_{abs}=100$  cm at  $\lambda=400$  nm



# SND counters assembling & installation



# Альфа Магнитный Спектрометр (AMS-02)



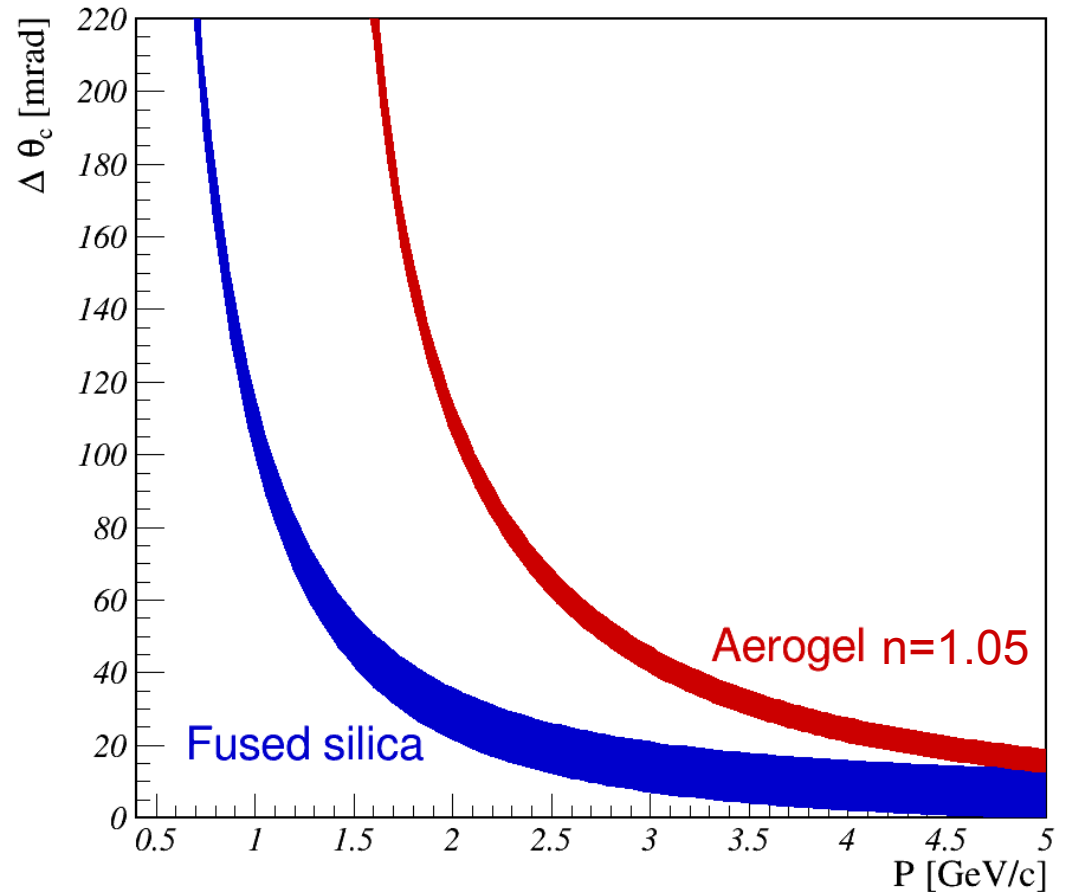
## Задачи:

- Антивещество
- Темная материя
- Состав космических лучей
- Поиск странглетов

- Работает на МКС с 2011 года
- $M = 8.5$  тон,  $V = 54$  м<sup>3</sup>,  $S = 1$  м<sup>2</sup>
- Магнитное поле – 1.26 кГс
- TOF, TRD, Si Tracker, RICH, Ecal

# Quartz vs Aerogel radiators

Difference in Cherenkov angle  
 $\theta_c$  for  $\pi$  and  $K$   
Bands – chromatic dispersion in  
350-700 nm



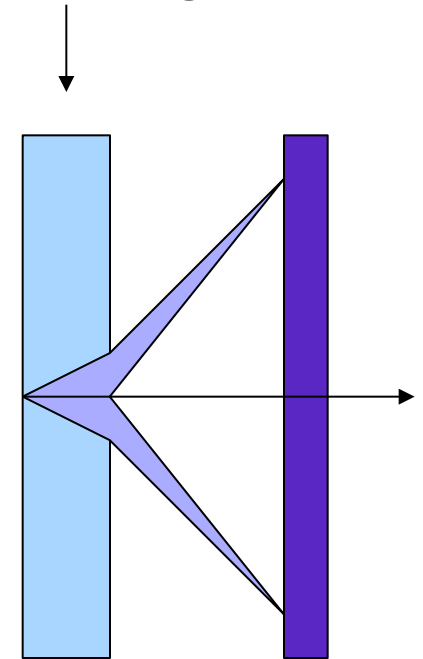
# FARICH perspectives

For the proximity focusing RICH detectors there are 3 main contributions to the

resolution:  $\sigma_{\Theta}^2 = \sigma_{\text{chr}}^2 + \sigma_{\text{geom}}^2 + \sigma_{\text{phot}}^2$

- Suggested technology of gradient aerogel tile production could give us radiators with  $\sigma_{\text{geom}} \ll \sigma_{\text{chr}}$
- Philips Digital Photon Counting are working on the next version of the sensor which could read out the time and micro-cell number (instead of the number of fired cells) of the hit,  $\sigma_{\text{phot}} \approx 20 \mu\text{m} \ll \sigma_{\text{chr}}$
- **Could we build RICH with  $\sigma_{\Theta}^2 \approx \sigma_{\text{chr}}^2$  ?**

Gradient aerogel tile



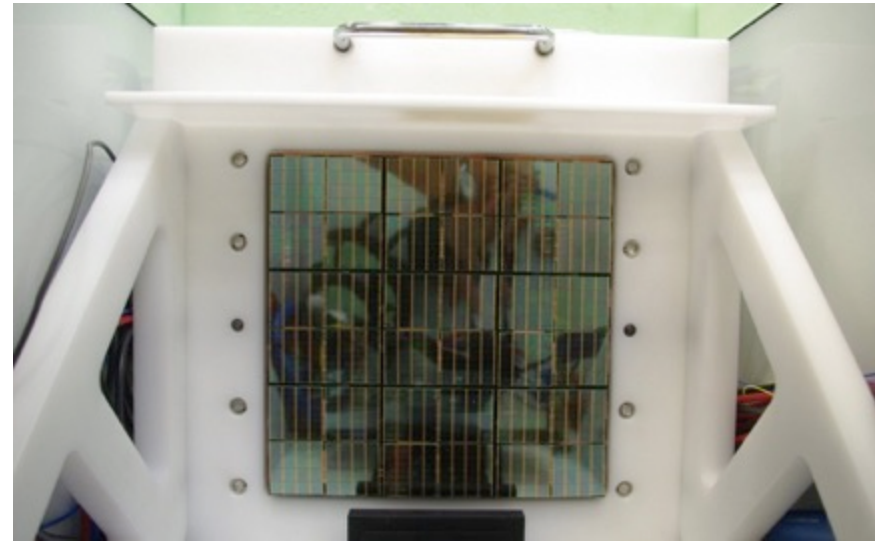
photon sensor with read out of the hit coordinate

# Beam test of FARICH at CERN PS T10, June 2012



## 4-layer aerogel

- $n_{\max} = 1.046$
- Thickness 37.5 mm
- Focal distance  
200 mm



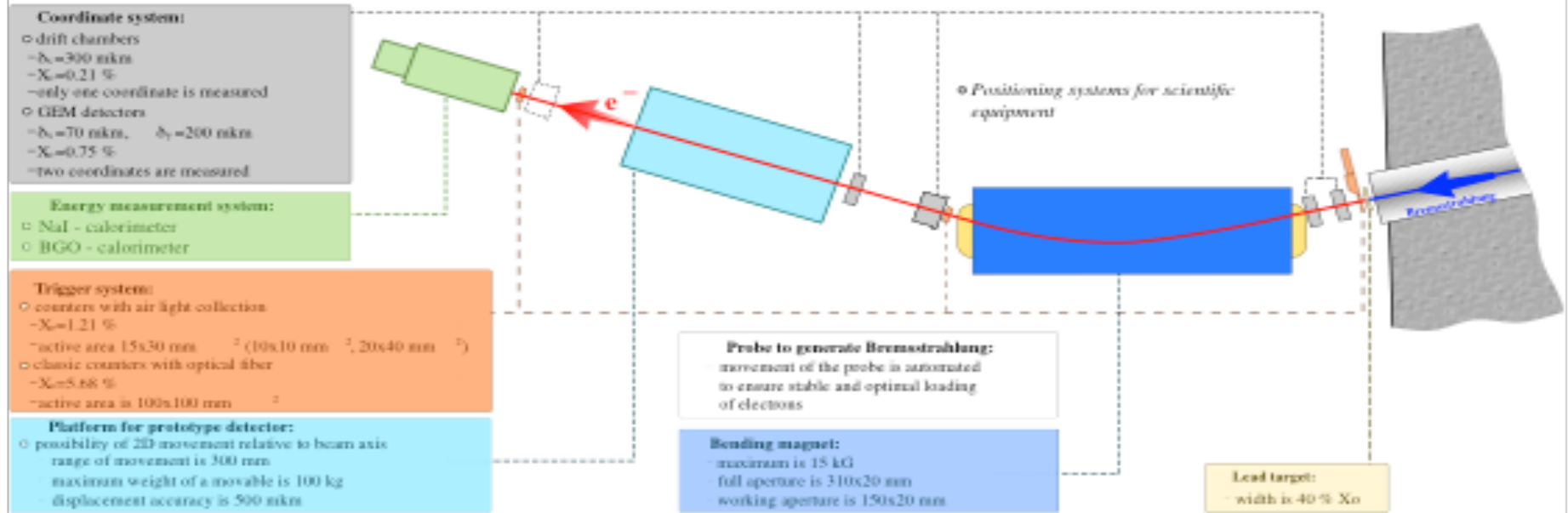
## Test conditions

- Positive polarity:  $e^+$ ,  $\mu^+$ ,  $\pi^+$ ,  $K^+$ ,  $p$
- Momentum: 1–6 GeV/c
- Trigger: a pair of sc. counters  $1.5 \times 1.5$   $\text{cm}^2$  in coincidence separated by  $\sim 3$  m
- No external tracking, particle ID, precise timing

## DPC matrix **$20 \times 20 \text{ cm}^2$**

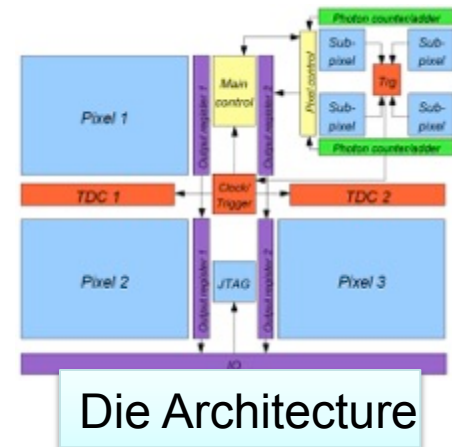
- Sensors: DPC3200-22-44
- 3x3 modules = 6x6 tiles =  
24x24 dies = 48x48 pixels
- 576 time channels
- 2304 amplitude (position) channels
- Operation at  $-40^\circ\text{C}$  to reduce  
dark counts

# Test beam facilities at BINP



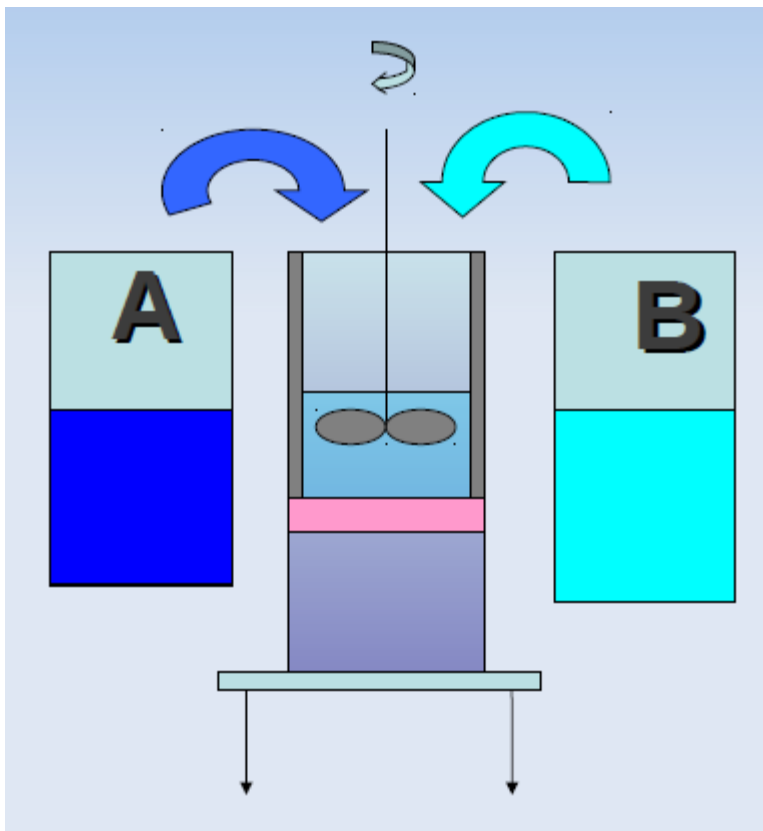
See V.S. Bobrovnikov et al., 2014 *JINST* 9 C08022 and G.N.Abramov et al., 2016 *JINST* 11 P03004

- Philips DPC matrix  $20 \times 20$   $\text{cm}^2$
- Sensors: DPC3200-22-44
- 3x3 modules = 6x6 tiles = 24x24 dies = 48x48 pixels
- 576 time channels
- 2304 amplitude (position) channels
- Size of pixel could be changed from  $3 \times 3$   $\text{mm}^2$  to  $50 \times 50$   $\mu\text{m}^2$
- Operation at  $-40^\circ\text{C}$  to reduce dark counts BNO-50, 2017





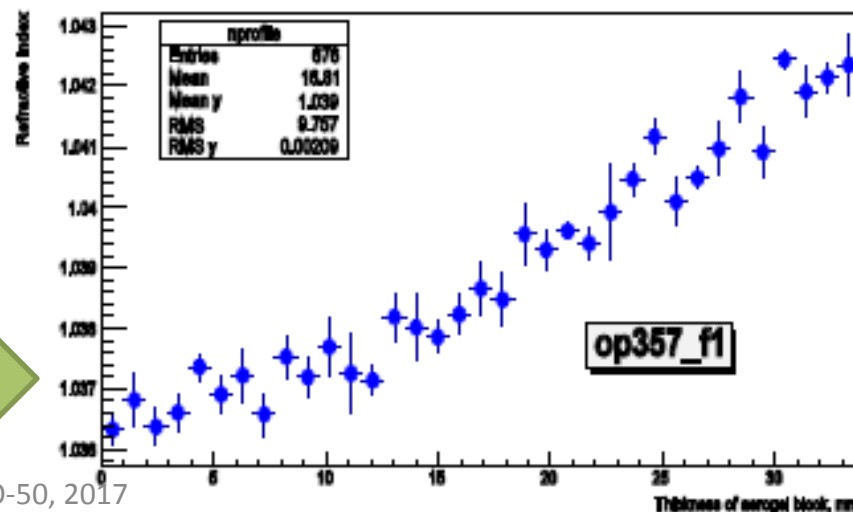
# Continuous density gradient aerogel



To produce aerogel tiles with designed profile of gradient we modernized the method suggested by [S.M. Jones “A method for producing gradient density aerogel”, J Sol-Gel Sci Technol. 44 (2007) 255]

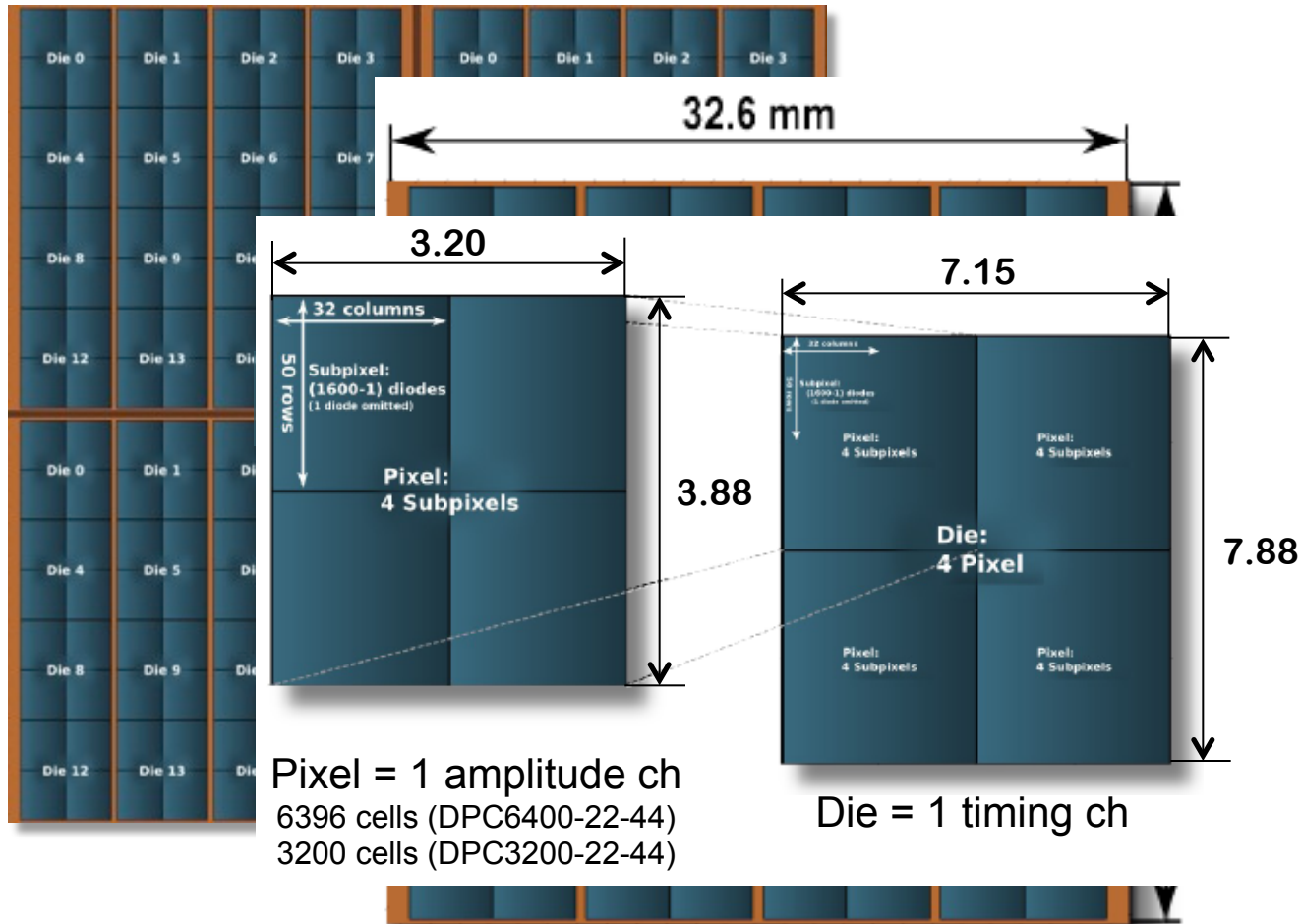
- We mix two pre-prepared mixtures with different content of TEOS fed by peristaltic pumps from vessels A and B.
- The mixture with designed concentration of TEOS seeps through the filter to the mould where gelation takes place.
- The mould is positioned on the vertically moving table. The peristaltic pumps and moving table are controlled by a computer.

Refr. Index vs X



Refractive index profile along thickness

# DPC hierarchy in PDPC-FARICH



Pixel-in-module packing density  $\approx 70\%$

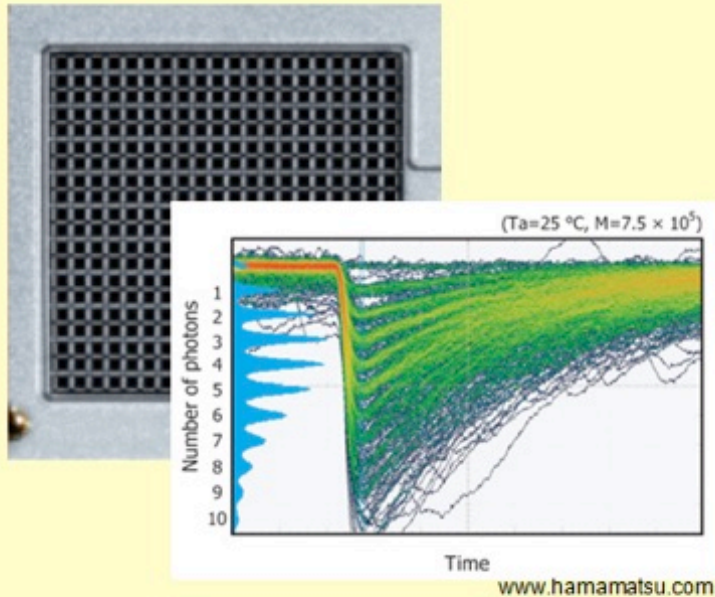




# DPC: Front-end Digitization by Integration of SPAD & CMOS Electronics

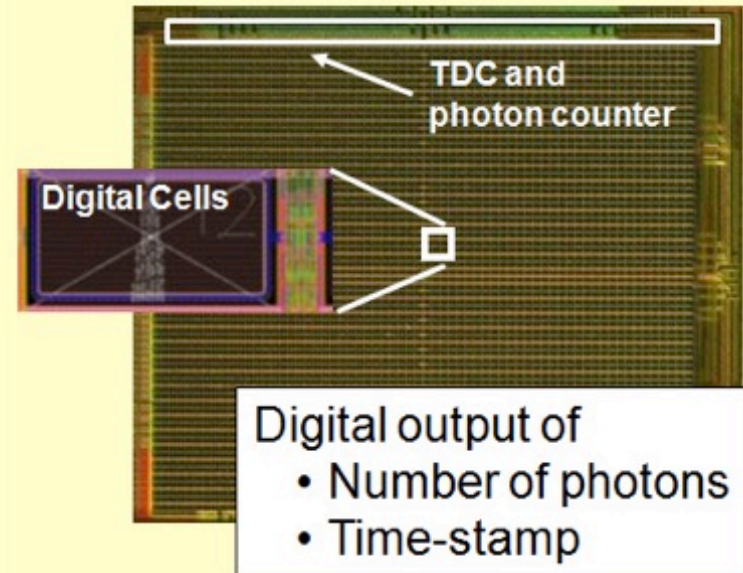
## Philips Digital Photon Counting (PDPC)

analog SiPM



Summing all cell outputs leads to an analog output signal and limited performance

Digital Photon Counter (DPC)



Integrated readout electronics is the key element to superior detector performance

T. Frach, G. Prescher, C. Degenhardt, B. Zwaans, IEEE NSS/MIC (2010) pp.1722-1727

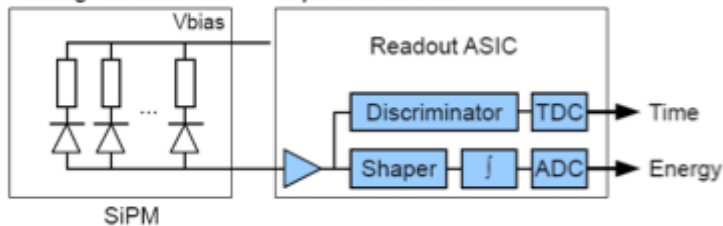
C. Degenhardt, T. Frach, B. Zwaans, R. de Gruyter, IEEE NSS/MIC (2010) pp.1954-1956

# DPC: Front-end Digitization by Integration of SPAD & CMOS Electronics

## Analog SiPM

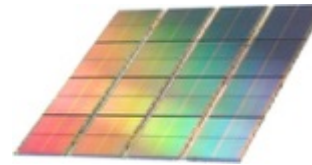


Analog Silicon Photomultiplier Detector

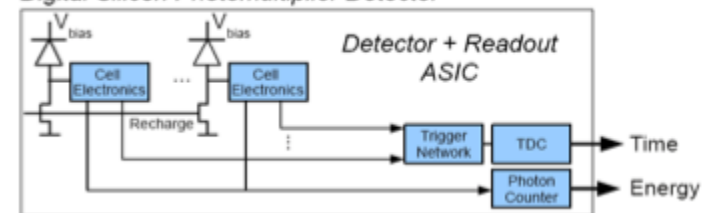


- discrete, limited integration
- analog signals to be digitized
- dedicated ASIC needed
- difficult to scale

## Digital SiPM



Digital Silicon Photomultiplier Detector



- fully integrated
- fully digital signals
- no ASIC needed
- fully scalable

# DPC tile – PCB with densely packed 4x4 sensors (8x8 pixels)

**DPC3200-22-44** – 3200 cells/pixel

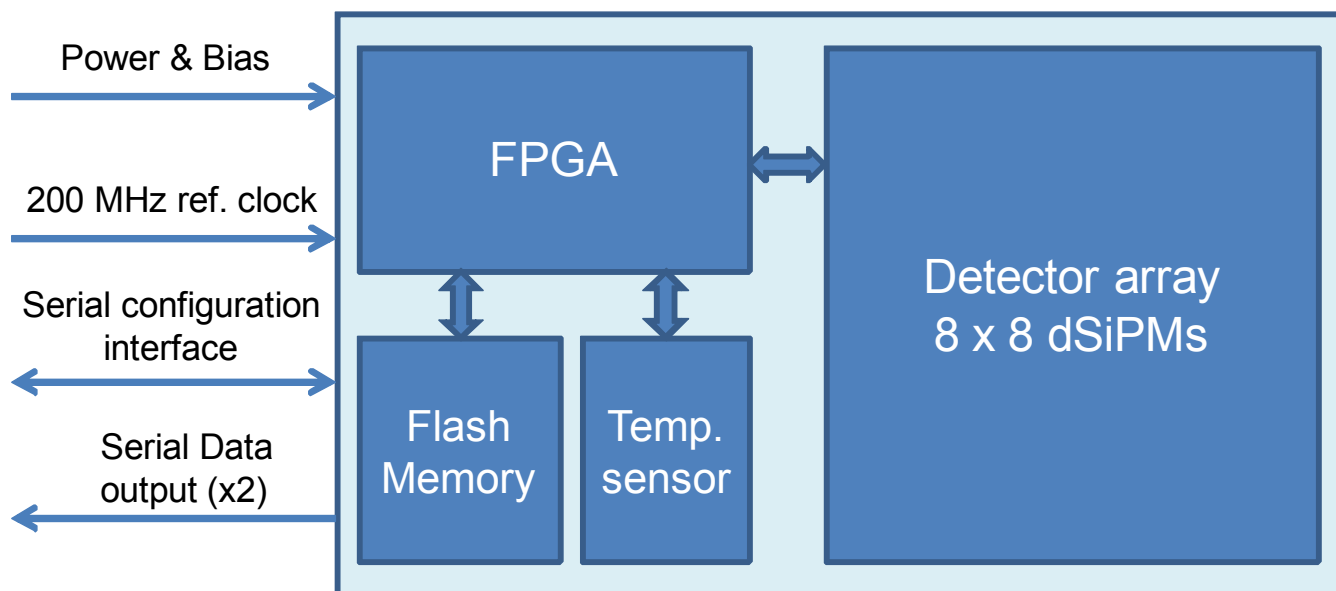
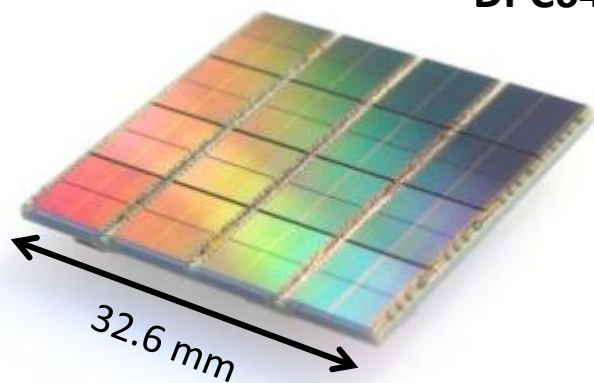
**DPC6400-22-44** – 6396 cells/pixel

## FPGA

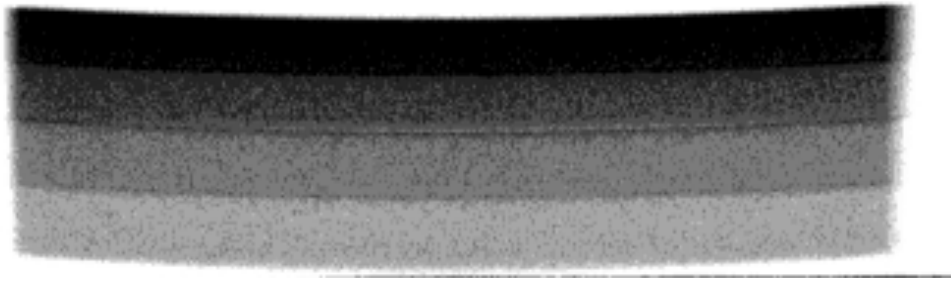
- Clock distribution
- Data collection/concentration
- TDC linearization
- Saturation correction
- Skew correction

## Flash

- FPGA firmware
- Configuration
- Inhibit memory maps



# Aerogel study with digital X-ray setup

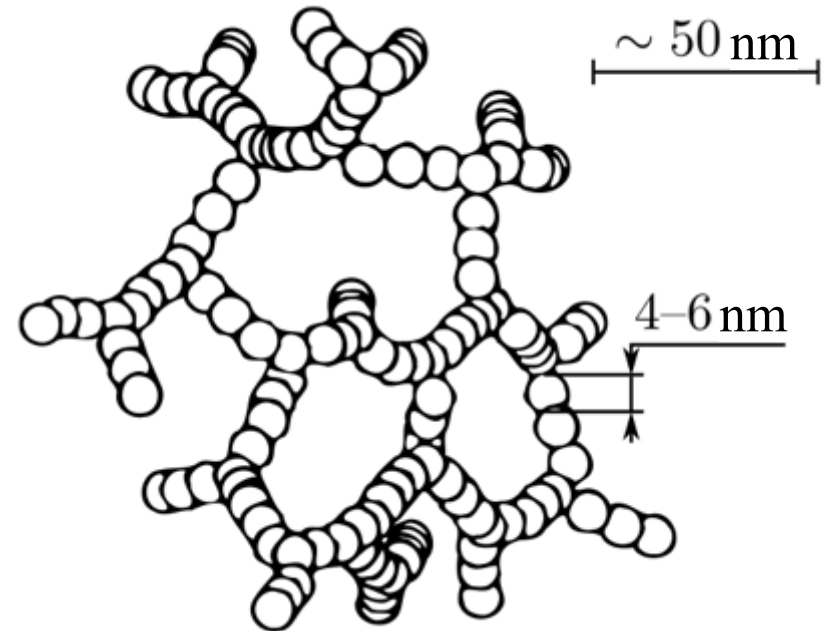
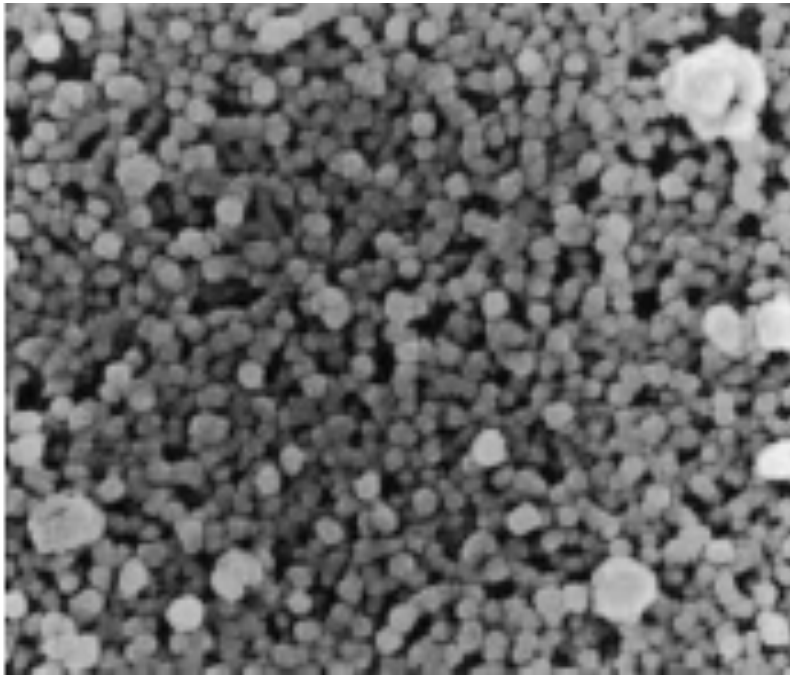


	n	h, mm
Layer 1	1.050	6.2
Layer 2	1.041	7.0
Layer 3	1.035	7.7
Layer 4	1.030	9.7



- $100 \times 100 \times 31 \text{ mm}^3$
- $L_{sc}(400\text{nm}) = 43 \text{ mm}$
- $n^2 = 1 + 0.438 * \rho$

# Refractive index



$\text{SiO}_2 + \text{H}_2\text{O}(1\div 5\%)$

$$n^2 = 1 + 0.438 \cdot \rho$$

$n=1.006\dots 1.070$  – synthesis

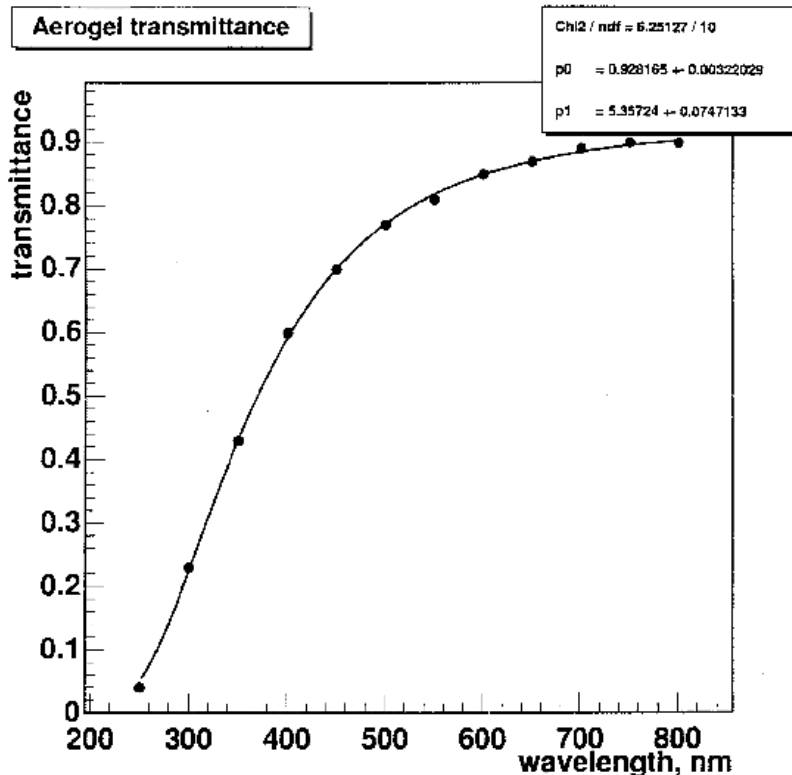
$n=1.070\dots 1.130$  – sintering

# Light scattering

Rayleigh scattering on aerogel structure elements

Transmittance:

$$T = \frac{I}{I_0} = A \cdot \exp \frac{-d}{L_{sc} \cdot (\lambda/400)^4} = A \cdot \exp \frac{-C \cdot d}{\lambda^4}$$



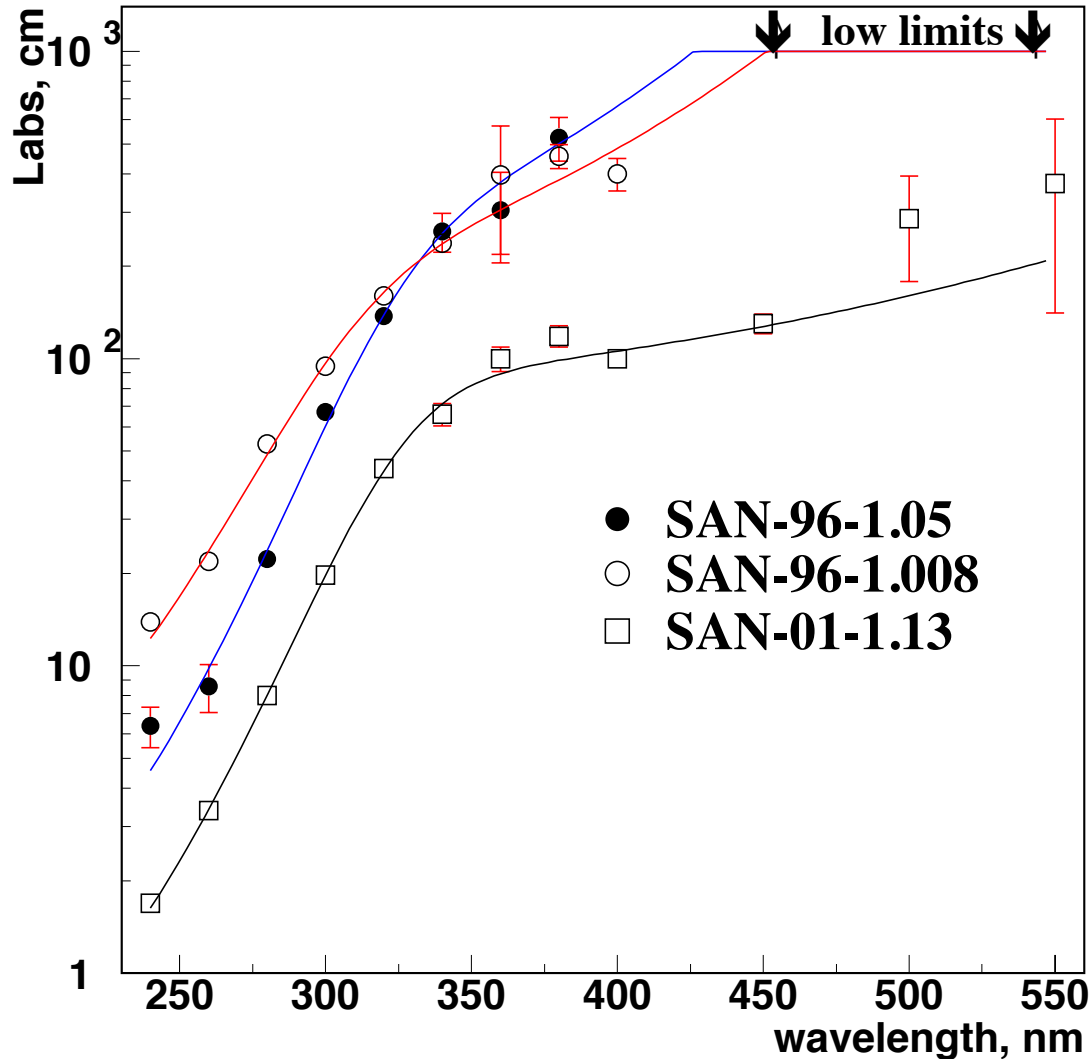
A – surface scattering coefficient  
 ~0.95 for intrinsic surface  
 ~0.70 for polished surface

$L_{sc}$  – scattering length at  $\lambda=400\text{nm}$   
 > 4.5 cm

C – clarity (  $0.4^4 / L_{sc}$  )  
 <  $0.0057 \mu\text{m}^4/\text{cm}$



# Light absorption

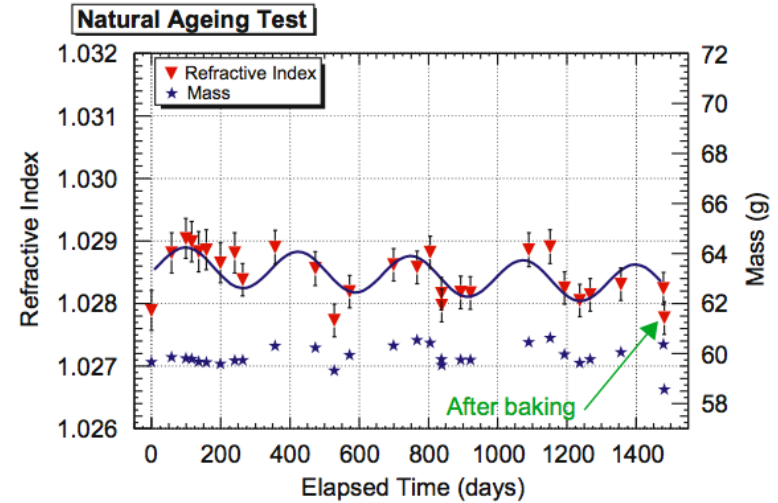
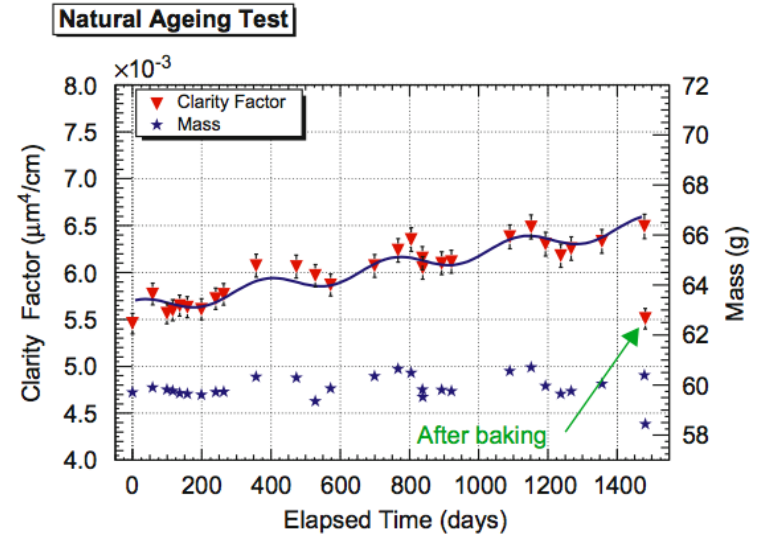
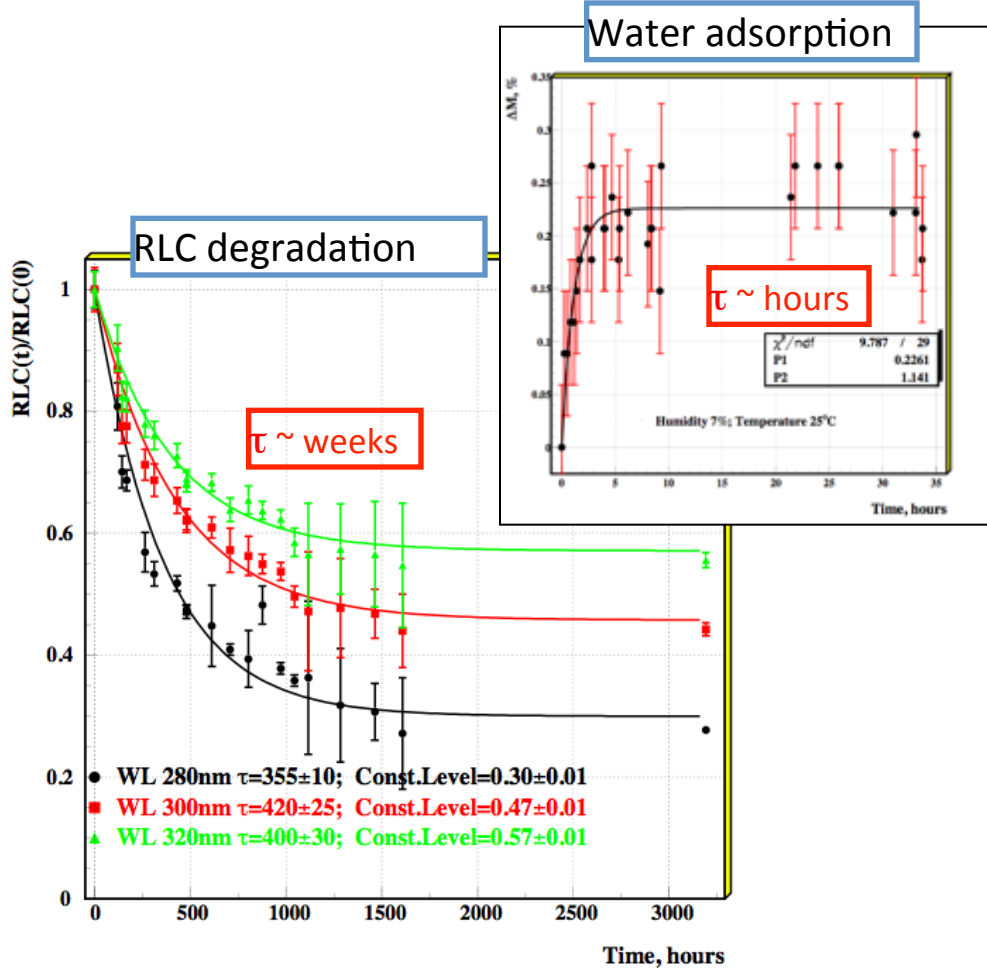


Light is absorbed by impurities.

Contamination of metals (Fe, Co, Cu, Mn, etc.) is determined by raw material quality and synthesis technology.

# Water adsorption

$$1 \text{ cm}^3 \Rightarrow S_{\text{inner}} \sim 100 \text{ m}^2$$

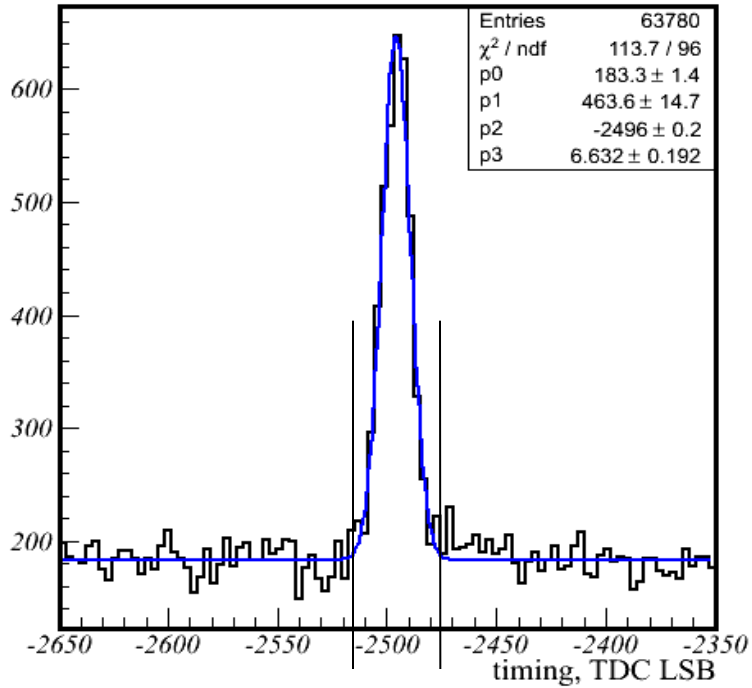


A.Yu.Barnyakov et al., NIM A598 (2009) 166

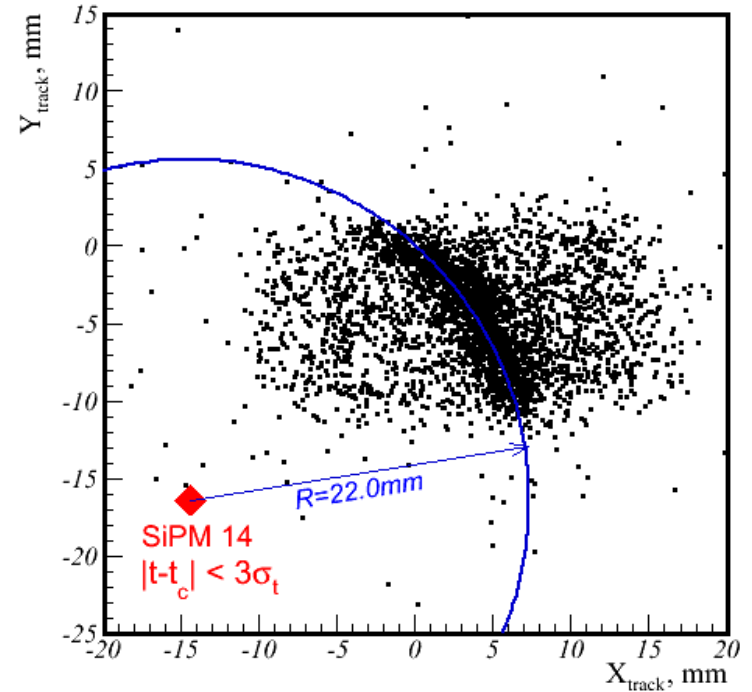
D.L.Perego, NIM A595 (2008) 224

# Event selection

Channel #14 phase adjusted timing

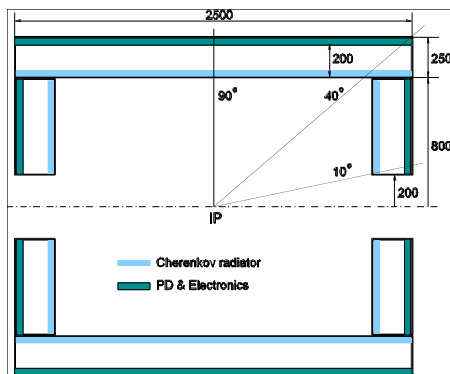


Hits in SiPM #14



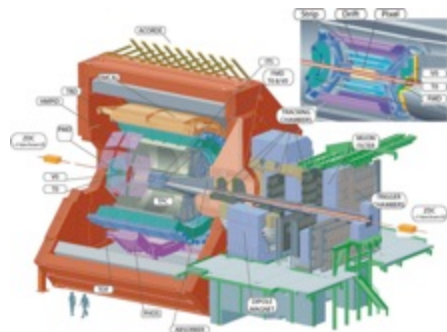
- We select events with  $|t-t_{\text{ch}}| < 3\sigma_t$

# FARICH proposals



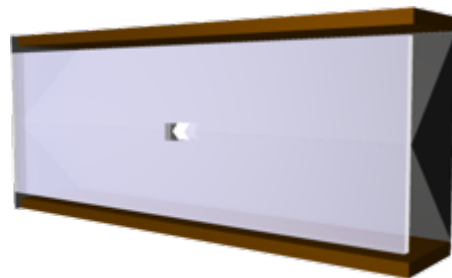
## FARICH for Super Charm-Tau Factory (Novosibirsk)

PID:  $\mu/\pi$  up to 1.7  $GeV/c$   
21m<sup>2</sup> detector area (SiPMs)  
~1M channels



## FARICH for ALICE HMPID upgrade

PID:  $\pi/K$  up to 10  $GeV/c$ ,  $K/p$  up to 15  $GeV/c$   
3m<sup>2</sup> detector area (SiPMs)



## Forward Spectrometer RICH for PANDA

PID:  $\pi/K/p$  up to 10  $GeV/c$   
3m<sup>2</sup> detector area (MaPMTs or SiPMs)